# Scintillation Detectors and Their Applications

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Graduate Research in Instrumentation and Detectors Summer School TRIUMF, UBC

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# Outline

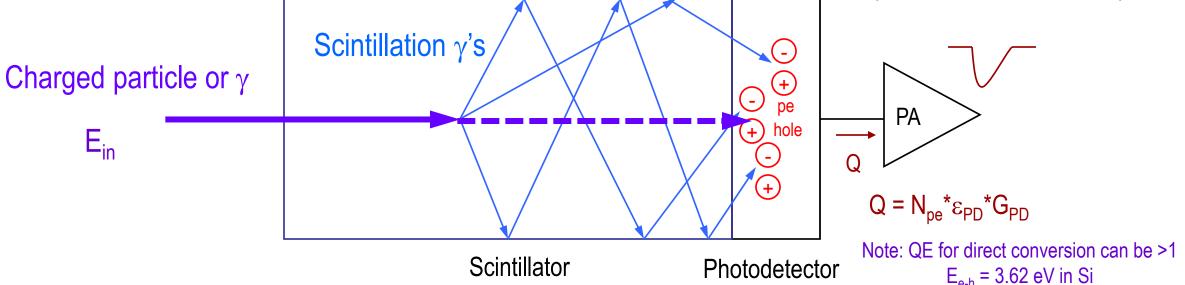
- Introduction
- Fundamentals of Scintillators
- Photodetectors
- Applications
  - Nuclear and Particle Physics
  - Astrophysics
  - Medical imaging
  - Nuclear Waste Management & Non-Proliferation

# **Scintillation Detectors**

- A scintillation detector is a device that detects the light produced by the deposition of energy via dE/dx in a scintillator. It consists of two basic elements:
  - Scintillator (produces the light)
  - Photodetector (detects the light)
- Can also include other elements such as additional absorbers which absorb part of the energy lost by a particle, light collectors and readout electronics.

### **Scintillation Detectors**

Incident particle can also interact directly in the photodetector to create e-h pairs



High energy particles deposit energy in the scintillator and produces scintillation photons

$$N_{scint} = E_{in} * \varepsilon_{con}$$

Scintillation photons are collected and detected in the photodetector producing *photoelectrons* 

$$N_{pe} = N_{scint} * \varepsilon_{col} * QE$$

# **Basic Types of Scintillators**

A scintillator is a material that transforms energy loss due to ionization (dE/dx) into light

### **Organic Scintillators**

- Plastics, liquids, organic crystals
- Low density (~ 1 g/cm<sup>3</sup>)
- Low Z
  - Requires interspersing high Z absorber material to achieve high stopping power for high energy γ's
  - n detection by (n,p) interactions
- Up to 10,000 γ/MeV
- ns decay times
- Relatively inexpensive
- Moderately rad hard (~ 10 kGy/yr)

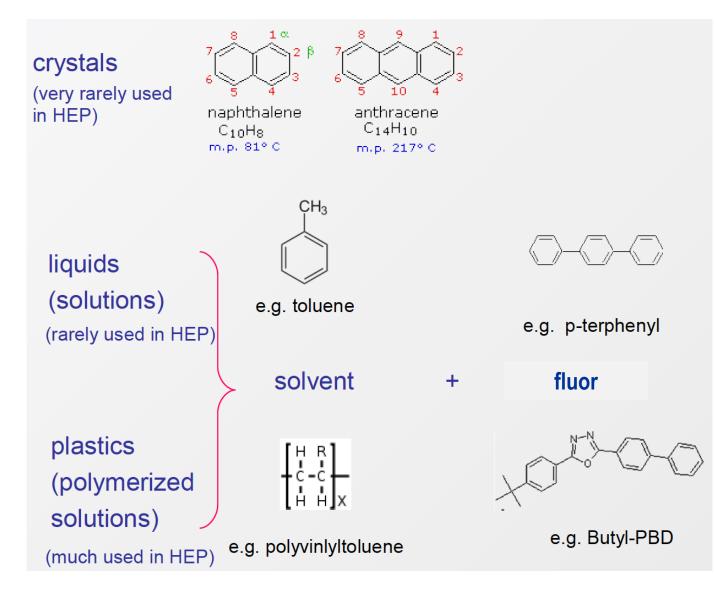
### **Inorganic Scintillators**

- Crystals
- Can high density (> 8 g/cm<sup>3</sup>)
- Typical have high Z
  - Leads to homogeneous detectors with very good energy resolution
  - Requires good light collection
  - Poor n detection efficiency
- Up to 50,000 γ/MeV
- ns to msec decay times
- Expensive
- Fairly radiation hard (~ 100 kGy/yr)

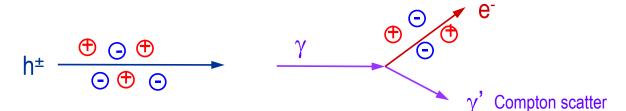
### **Noble Liquids Scintillators**

- LAr, LKr, LXe
- Requires working at cryogenic temperatures
- Moderate density and Z
  - (~ 1.4- 3.0 g/cm<sup>3</sup>)
- Scintillation in UV or VUV
- Light yield ~ 50,000 γ/MeV
- Produce ionization plus scintillation

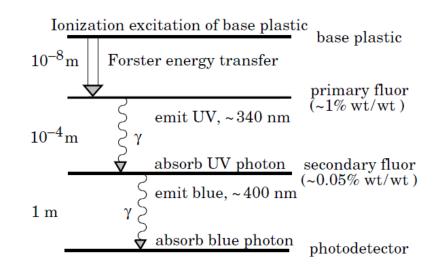
### **Organic Scintillators**



### **Plastic Scintillators**



- Charged particles produce ionization which causes excitations in the medium
- Gammas interact and produce electrons which cause ionization
- Primary ionization causes excitation of molecules in the base plastic
- De-excitation of the base plastic produces scintillation photons in the short wavelength UV (~ 300 nm), which are absorbed by a primary fluor, or energy is transferred directly from the base to the fluor in a very short distance.
- Primary fluor emits at a longer wavelength (~ 340 nm), which is absorbed by a secondary fluor that emits in the visible (~ 400 nm)
- Each process involves losses due to the efficiency of energy transfer at each step, resulting in an overall light yield of ~ 1 photon per 100 eV of energy deposit



Review of Particle Physics Journal of Physics G, Vol 37, No 7A (2010)

### **Wavelength Shifters**

Wavelength shifters absorb light at a short wavelength and re-emit at a longer wavelength.

Change in wavelength between absorption and re-emission is called the Stokes Shift

Wavelength shifting can take place in the base plastic scintillator, or can be applied to the light produced by the scintillator using other materials such as wavelength shifting bars or fibers.

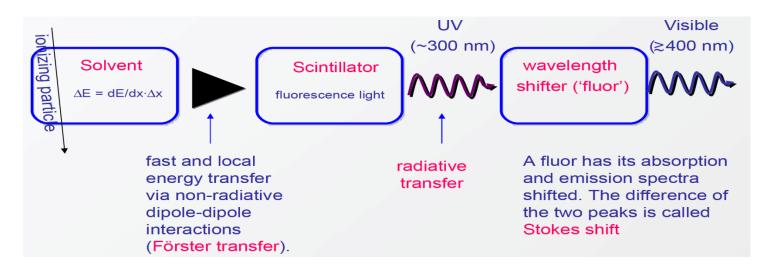
#### Wavelength shifting can accomplish several goals:

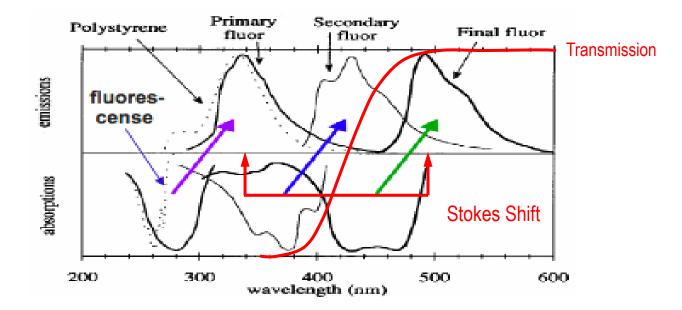
- 1. Improve light propagation in the material
- Attenuation length of most base materials is much longer at longer wavelengths
- 2. Redirect the light for better light collection
- 3. Provide a better wavelength match to the photodetector

#### These features come at some cost:

- 1. Efficiency of wavelength shifting may be low
- 2. Primary light must be collected by the wavelength shifter
- 3. Decay time of wavelength shifter may be slower than primary emitter

### **Wavelength Shifters**





### **Commonly Used Plastic Scintillators**

Scintillator	Light output (% Anthracene)	Decay time (ns)	Wavelength of emission max (nm)	Attenuation length (cm)	Application
BC 400	65	2.4	423	250	General purpose
BC 404	68	1.8	408	120	Fast counting
BC 408	64	2.1	425	380	Large area
BC 412	60	3.3	434	400	Large area
BC 418	67	1.4	391	100	Ultrafast timing
BC 428	50	12.0	490	330	Green emitting

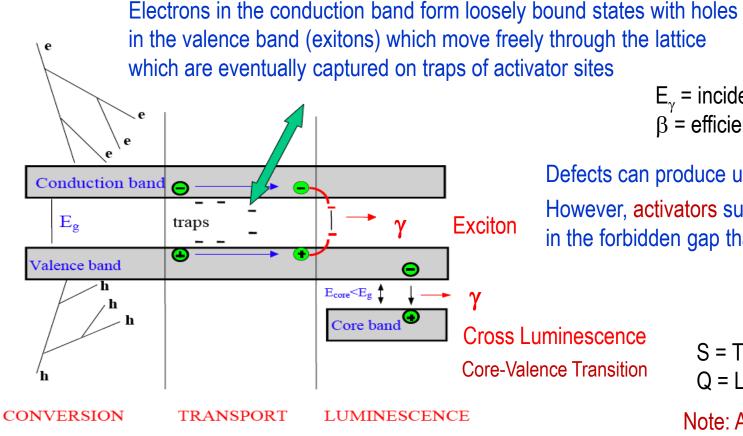
**Bicron/Saint-Gobain** 

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#### Light output of Anthracene ~ 40,000 $\gamma$ /MeV (similar to NaI)

### **Crystal Scintillators**

Primary ionization excites an electron from the valence band to the conduction band leaving a hole in the valence band



$$n_{e-h} = \frac{E_{\gamma}}{\beta E_{gap}}$$

 $E_{\gamma}$  = incident gamma ray energy  $\beta$  = efficiency factor for e-h pair formation ( $\beta \sim 2-4$ )

Defects can produce unwanted traps in the forbidden gap However, activators such as TI and Ce can create energy levels in the forbidden gap that can produce luminescence

$$N_{photon} = n_{e-h} \cdot S \cdot Q$$

S = Transfer efficiency to luminescence center Q = Luminescence efficiency

Note: Activator determines wavelength and decay time of emission

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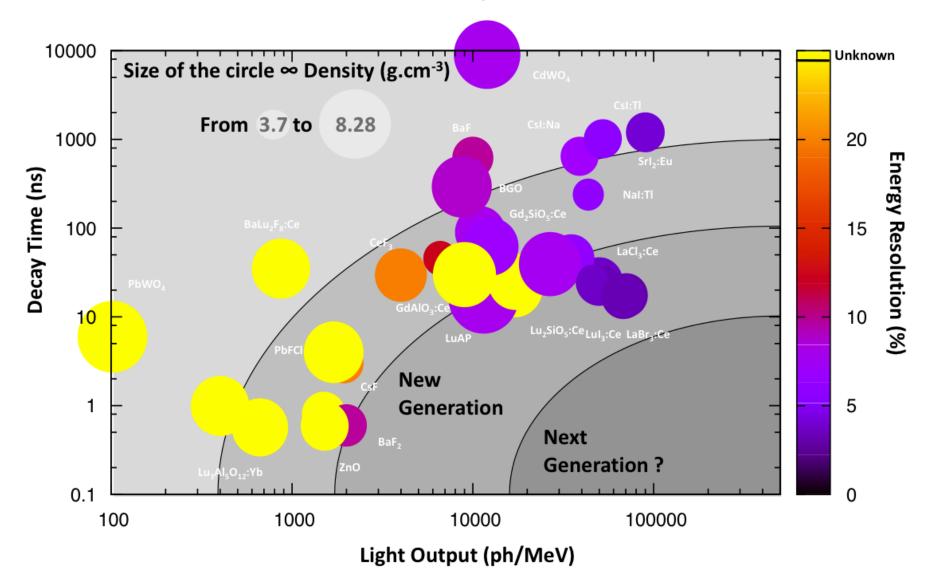
### **Commonly Used Crystal Scintilators**

Parameter	r: ρ	MP	$X_0^*$	$R_M^*$	$dE^*/dx$	$\lambda_I^*$	$\tau_{ m decay}$	$\lambda_{ m max}$	$n^{ atural}$			d(LY)/dT
Units:	$g/cm^3$	°C	$\mathbf{cm}$	$\mathrm{cm}$	MeV/cm	$\mathbf{cm}$	$\mathbf{ns}$	nm		output <sup>†</sup>	scopic?	$\%/^{\circ}C^{\ddagger}$
NaI(Tl)	3.67	651	2.59	4.13	4.8	42.9	230	410	1.85	100	yes	-0.2
BGO	7.13	1050	1.12	2.23	9.0	22.8	300	480	2.15	21	no	-0.9
$BaF_2$	4.89	1280	2.03	3.10	6.5	30.7	$630^{s}$	$300^s$	1.50	$36^s$	no	$-1.3^{s}$
							$0.9^{f}$	$220^{f}$		$3.4^{f}$		${\sim}0^{f}$
$\operatorname{CsI}(\operatorname{Tl})$	4.51	621	1.86	3.57	5.6	39.3	1300	560	1.79	165	$\operatorname{slight}$	0.3
CsI(pure)	4.51	621	1.86	3.57	5.6	39.3	$35^{s}$	$420^{s}$	1.95	$3.6^{s}$	slight	-1.3
							$6^{f}$	$310^{f}$		$1.1^{f}$		
$PbWO_4$	8.3	1123	0.89	2.00	10.1	20.7	$30^s$	$425^{s}$	2.20	$0.083^{s}$	no	-2.7
							$10^{f}$	$420^{f}$		$0.29^{f}$		
$\mathrm{LSO(Ce)}$	7.40	2050	1.14	2.07	9.6	20.9	40	402	1.82	83	no	-0.2
LaBr <sub>3</sub> (Ce	) 5.29	788	1.88	2.85	6.9	30.4	20	356	1.9	130	yes	0.2

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Light output of NaI(TI) ~ 40,000  $\gamma$ /MeV

### **Characteristicsof Crystal Scintillators**

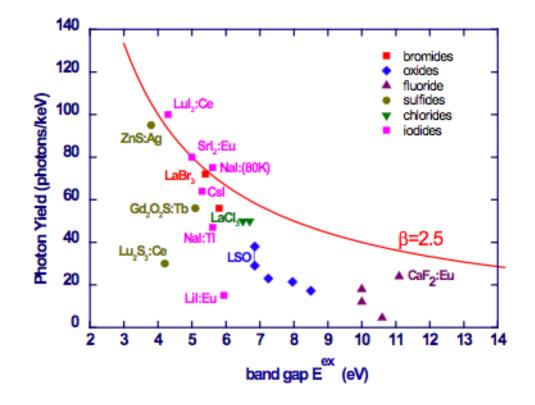


### **Crystal Scintillators**



### **Energy Resolution in Scintillators**

$$(\sigma_{\rm E}/{\rm E})^2 = (1/{\rm N}_{\rm pe}) + (\sigma_{\rm int}/{\rm E})^2 + (\sigma_{\rm np}/{\rm E})^2$$



$$N_{ph} \le N_{eh} = \frac{E_{\gamma}}{\beta E_{gap}}$$

 $\sigma_{int}$  = Intrinsic Non-Uniformity  $\sigma_{np}$  = Non-Proportionality

### **Scintillator Non-Proportionality**

Plastic scintillators do not respond linearly to ionization density

Saturation effects occur due to quenching and recombination

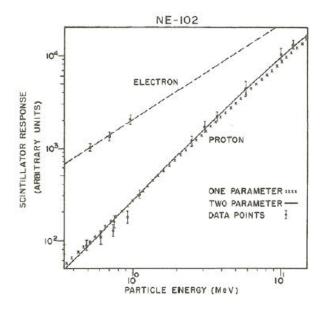
 $dE/dx \sim 2 MeV/cm$  for minimum ionizing particles

Semi-emperical formula:

Birks Law

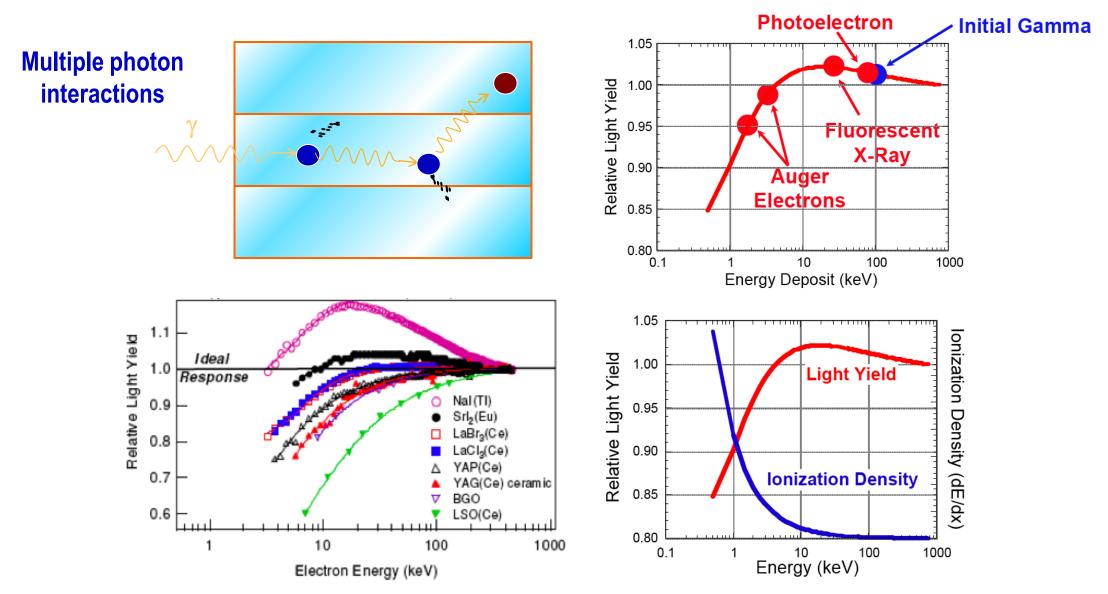
$$\frac{dL}{dx} = L_0 \frac{dE / dx}{1 + k_B dE / dx}$$

*k*<sub>B</sub> = Birks' constant determined by measurement



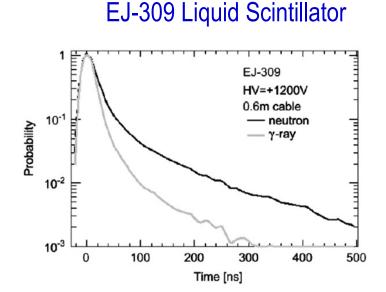
G.Knoll, Radiation Detection and Measurement

# **Non-Proportionality in Crystals**

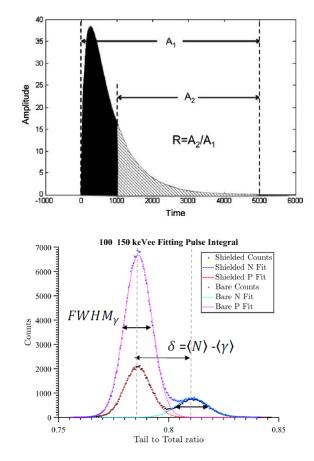


### **Pulse Shape Discrimination**

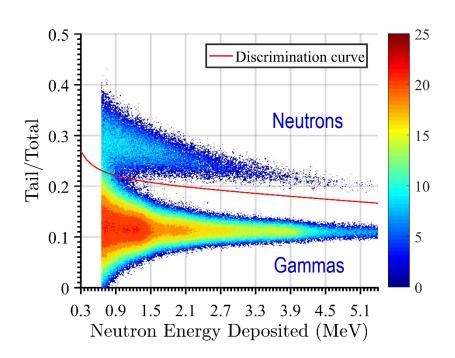
Non-Proportionality in Organic Scintillators can be used for Pulse Shape Discrimination



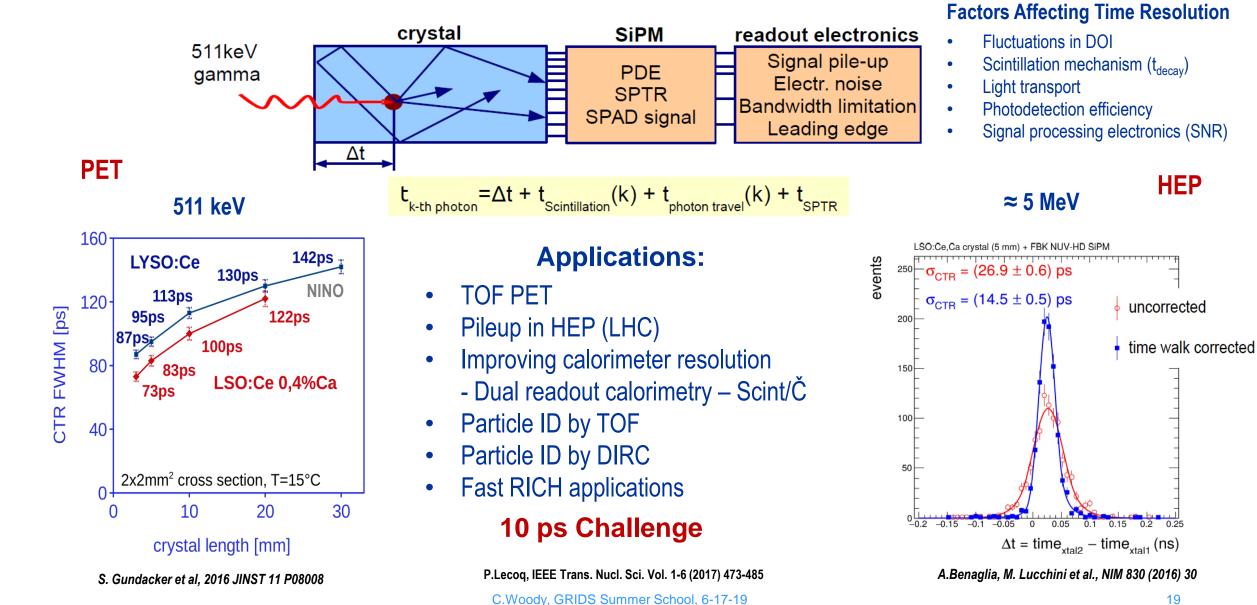
Fast scintillation is quenched Slow Fluorescence not quenched



#### Tail/Total Ratio provides n-y separation



# **Fast Timing with Scintillators**



# **Radiation Damage**

Radiation can produce defects in a scintillating material resulting in a loss in transmission and/or scintillation light output

### **Ionization Damage**

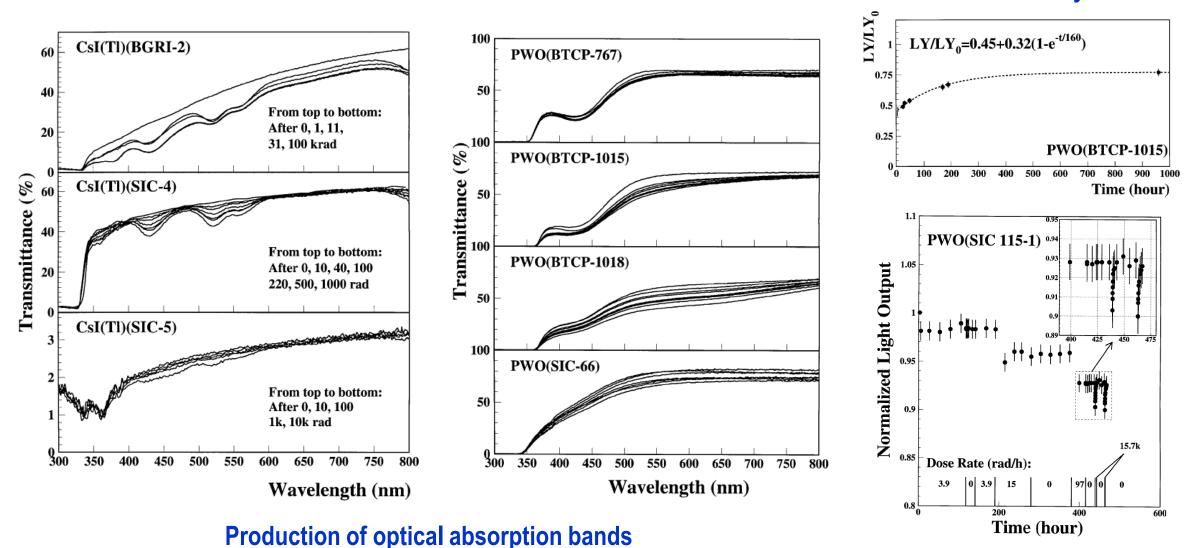
- Electrons are liberated and are captured in existing traps in the crystal or plastic
- These traps can then cause resonance absorption centers for the emitted scintillation photons which often appear as optical absorption bands
- Shallow resonance traps can often be annealed away by thermal or optical annealing

### **Displacement Damage**

- Heavy particles (charged hadrons, neutrons, heavy ions, etc..) can cause displacement damage in the material and produce additional defects
- These defects can be permanent (e.g., due to nuclear breakup) or are harder to anneal (interstitial-vacancy recombination)
- However, generally, the scintillation mechanism is not damaged

# **Radiation Damage in Crystals**

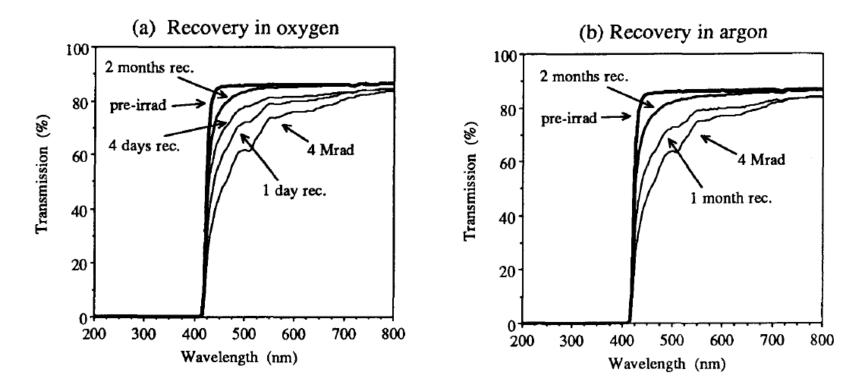
Recovery



**Dose rate dependence** 

R.Y.Zhu, Nucl. Inst. Meth. A413 (1998) 297-311

### **Radiation Damage in Plastics**

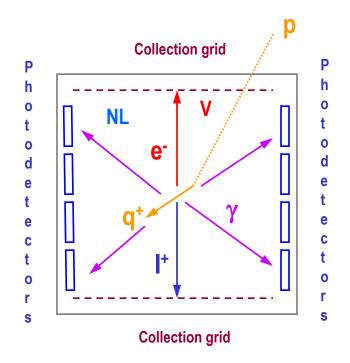


<u>Figure 4</u>: Transmission spectra of cylindrical BC408 samples (15 mm diameter x 10 mm thickness) after being irradiated (in air) to 4 Mrad by a  $^{60}$ Co source (140 krad/hr) and allowed to recover in flowing (a) oxygen and (b) argon atmospheres.

#### Annealing in oxygen is more effective due to reduction of oxygen vacancies

C.Zorn, IEEE Trans. Nucl. Sci . Vol. 37-2 (1990) 504-512

# **Noble Liquid Scintillators**



- Moderate density
- High scintillation light yield
- Emission in the VUV
- Possibility of measuring scintillation light and collecting charge
- Drift time can give position (TPC)

	LAr	LKr	LXe	
Atomic Number (Z)	18	36	54	
Density (g/cm3)	1.40	2.41	2.95	
Boiling Point (deg C @ 1atm)	87.3	119.8	165.0	
Radiation length (cm)	14.0	4.7	2.4	
Moliere radius (cm)	8.0	5.5	4.2	
Interaction Length (cm)	84	61	57	
Luminescence (nm)	128	147	174	
Light yield	102	~100	116	
Decay time (ns) fast	6.5	2.0	2.2	
slow	1100.0	85.0	27.0	

Some values from R.Wigmans, Calorimetry, Energy Measurement in Particle Physics

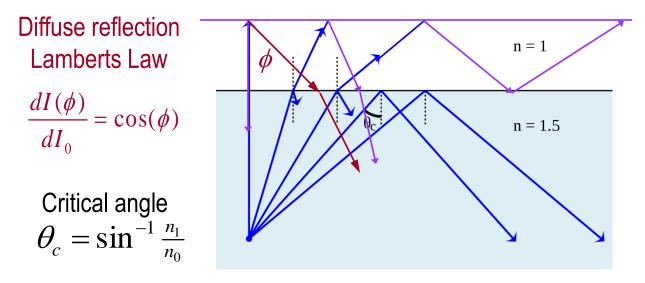
# **Light Collection**

Light produced in the gamma converter must be collected and directed onto the photodetector.

### Light collection techniques

- Direct reflection or use of reflective materials
- Light guides
- Wavelength shifters
- Factors to consider
- Efficiency
- Uniformity
- Geometry
- Matching to readout device

### Reflectors



Best reflector is Total Internal Reflection

- Scintillation light is produced isotropically
- Some fraction F will be trapped in the scintillator by internal reflection
  - requires good polished surface
  - however, may lead to non-uniformities due to self absorption
- Escaping fraction can be reflected back into scintillator by an external reflector
  - Requires an air gap between scintillator surface and reflector to preserve internal reflection
  - Specular reflector preserves angle
  - Diffuse reflector can change the angle of exiting rays to improve efficiency

### **Focusing Light Guides**

Need to concentrate the light, but Phase Space density  $\rho(q,p)$  remains constant with time

Louisville's Theorem

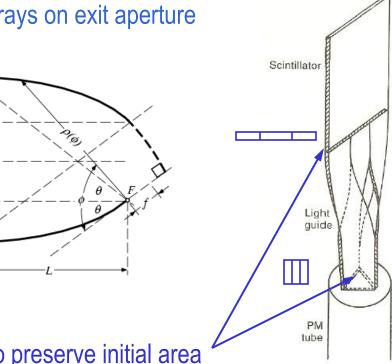
 $\frac{d\rho}{dt} = \sum_{i=1}^{n} \left( \frac{\partial\rho}{\partial q_{i}} \dot{q}_{i} + \frac{\partial\rho}{\partial p_{i}} \dot{p}_{i} \right) = 0$ 

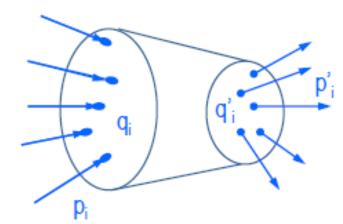


Winston Cone

Adiabatic Light Guide

> G.Knoll, Radiation Detection and Measurement

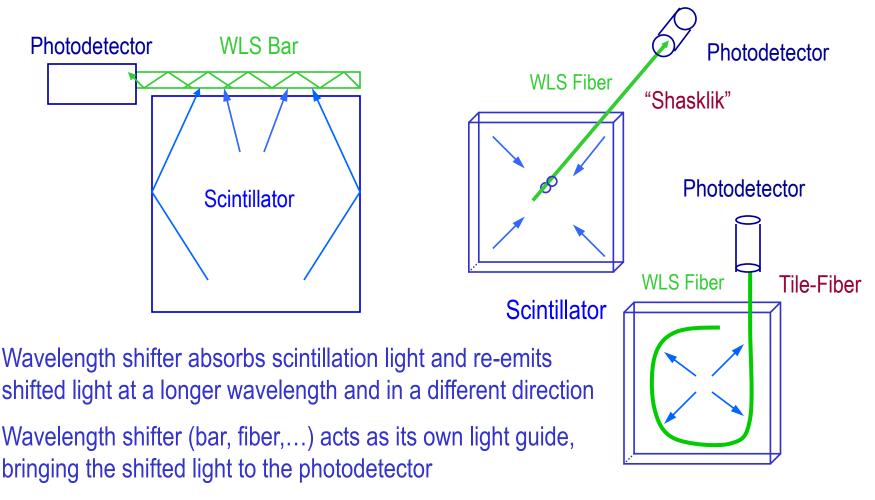




Reducing area increases divergence

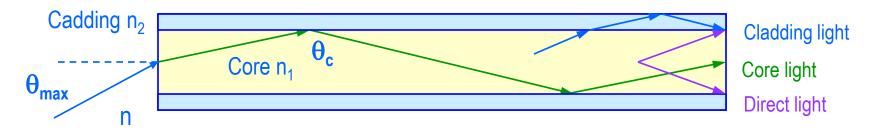
Try to preserve initial area

### **Wavelength Shifters as Light Collectors**



Note: Transfer of light from scintillator to wavelength shifter, absorption, re-emission, collection and transfer to photodetector can be very inefficient

### **Optical Fibers**



Fibers are used both as light guides and as active detector material - Scintillator or Cherenkov

Light is trapped by internal reflection between core and cladding  $(n_1 > n_2)$ Numerical Aperture defines the maximum angle that can be trapped

$$N.A. = n \cdot \sin \theta_{\max} = \sqrt{n_1^2 - n_2^2}$$

Trapping fraction is small

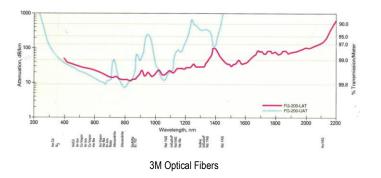
 $\mathbf{F}_{fib} = \frac{1}{2} \left( 1 - \frac{n_2}{n_1} \right)$ 

G.Knoll, Radiation Detection and Measurement

 $n_1$ =1.58 (polystyrene)  $n_2$ =1.49 (PMMA)  $F_{fib} \sim 3\%$ x2 ~ 6% if use reflector or read out both ends

Attenuation length is highly wavelength dependent Some escaping light is trapped in the cladding Direct light can also escape

All can lead to non-uniformities in Sci-Fi applications



### **Photodetectors**

Must convert the light from the scintillator into electrons which can then be amplified and measured

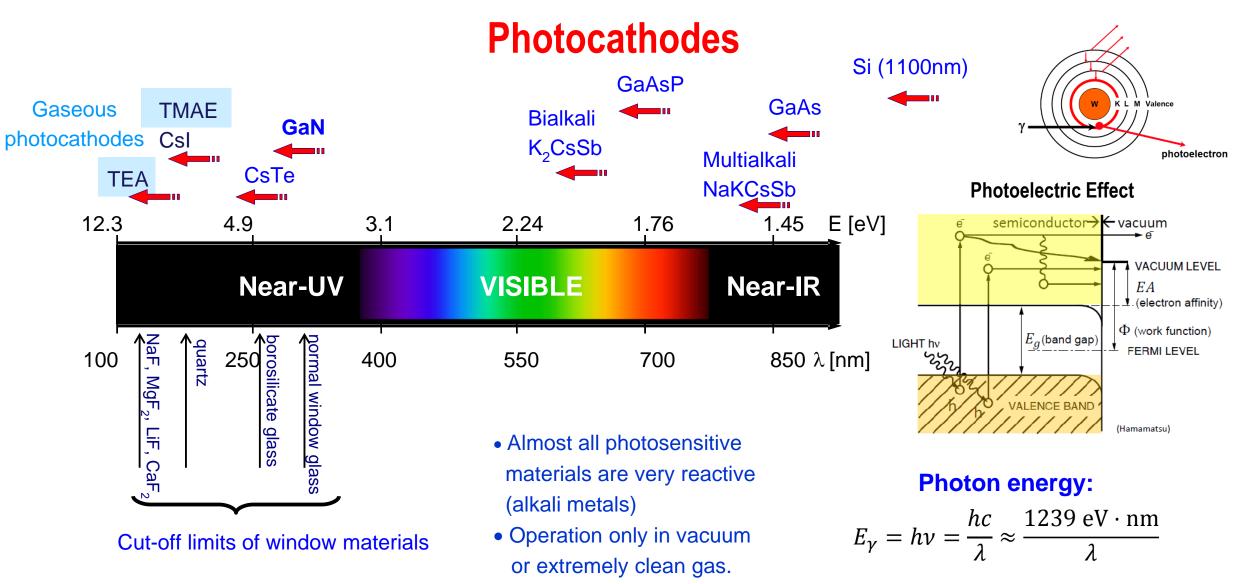
Photodetectors consist of two basic elements:

Photocathode

- Converts photons into photoelectrons via the photoelectric effect
- Performance determined by the Quantum Efficiency

### Charge amplification

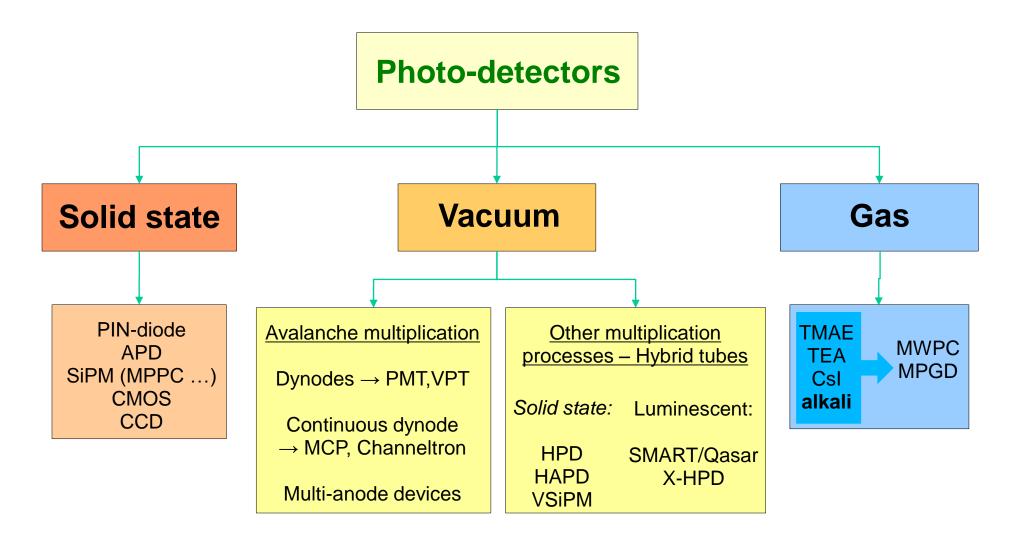
- Can involve gain or only direct conversion
- Performance determined by signal to noise ratio
- Requires appropriate readout electronics



Visible range 400 nm – 780 nm  $\rightarrow$  3.1 - 1.6 eV

Exceptions: Csl, Si

# **Types of Photodetectors**



### **Properties of Various Photodetectors**

Type	$\lambda$ (nm)	$\epsilon_Q \epsilon_C$	Gain	Risetime (ns)	$\begin{array}{c} {\rm Area} \\ ({\rm mm}^2) \end{array}$	1-p.e noise (Hz)	HV (V)	Price (USD)			
PMT* MCP* HPD* GPM* APD PPD	$\begin{array}{c} 115-1700\\ 100-650\\ 115-850\\ 115-500\\ 300-1700\\ 320-900 \end{array}$	$\begin{array}{c} 0.15 – 0.25 \\ 0.01 – 0.10 \\ 0.1 – 0.3 \\ 0.15 – 0.3 \\ \sim 0.7 \\ 0.15 – 0.3 \end{array}$	$\begin{array}{c} 10^3 - 10^7 \\ 10^3 - 10^7 \\ 10^3 - 10^4 \\ 10^3 - 10^6 \\ 10 - 10^8 \\ 10^5 - 10^6 \end{array}$	$0.7{-}10$ $0.15{-}0.3$ 7 O(0.1) O(1) $\sim 1$	$\begin{array}{c} 10^2 - 10^5 \\ 10^2 - 10^4 \\ 10^2 - 10^5 \\ O(10) \\ 10 - 10^3 \\ 1 - 10 \end{array}$	$\begin{array}{c} 10 - 10^4 \\ 0.1 - 200 \\ 10 - 10^3 \\ 10 - 10^3 \\ 1 - 10^3 \\ O(10^6) \end{array}$	$\begin{array}{c} 500{-}3000\\ 500{-}3500\\ {\sim}2\times10^4\\ 300{-}2000\\ 400{-}1400\\ 30{-}60\end{array}$	$\begin{array}{c} 100{-}5000\\ 10{-}6000\\ \sim 600\\ O(10)\\ O(100)\\ O(100)\\ O(100)\end{array}$			
							Deview of Destinte Device				

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- **PMT** = Photomultiplier tube (workhorse, but doesn't work in magnetic fields)
- **MCP** = Multichannel Plate Detector (dense, small diameter charge amplifiers fast)
- **HPD** = Hybrid Photodiode Detector (vacuum PMT with silicon sensor)
- **GPM** = Gaseous Photon Detector (solid & gaseous photocathode can cover large areas)
- APD = Avalanche Photodiode (also PIN diode)
- **PPD** = Pixellated Photon Detector (SiPM, GPMT, MPPC,...)

### **Photomultiplier Tubes**

#### **Principle of operation:**

- photo-emission from the photocathode -QE
- collection of photoelectrons by 1<sup>st</sup> dynode
- Secondary emission (SE) from N dynodes:
  - dynode gain  $\delta_i \sim 3 50$  (function of incoming electron energy);
  - total gain *M*:

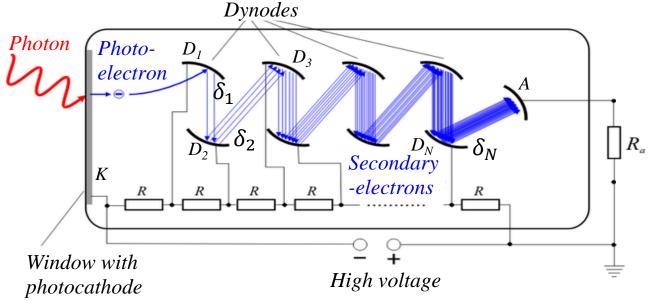
$$\mathbf{M} = \delta_1 \cdot \delta_2 \cdots \delta_N = \prod_{i=1}^N \delta_i$$

- Example:
  - 10 dynodes with
    - $\delta = 4$
    - $M=\delta^N=4^{10}\approx 10^6$





The Workhorse

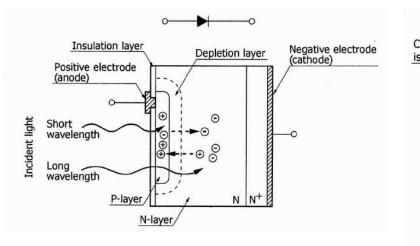


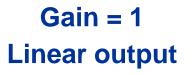
# **Silicon Photodetectors**

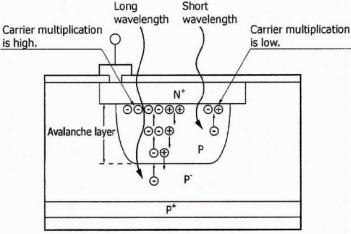
#### **PIN Diodes**

### **Avalanche Photodiodes (APDs)**

### **Geiger APDs (GAPDs)**







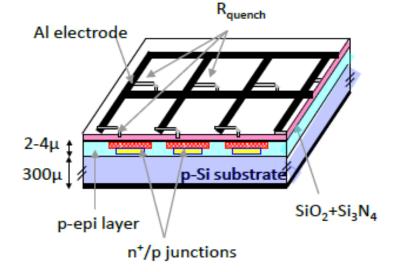
Gain = ~ 50-500 V<sub>bias</sub> < V<sub>breakdown</sub> Linear output cathode anode n+ n-implant n+ \* substrate (a)

Gain = ~ 50-500 V<sub>bias</sub> > V<sub>breakdown</sub> "Quantized" output

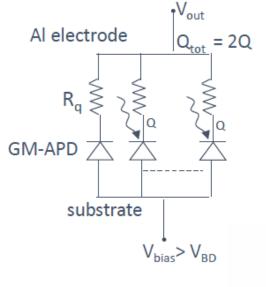
# Silicon Photomultipliers (SiPMs)

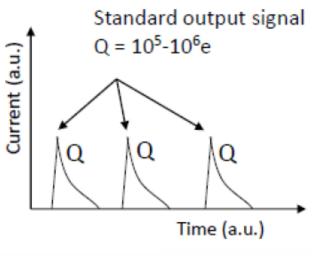
Silicon Photomultipliers are arrays of Single Photon Avalanche Diiodes (SPADs) that are biased slightly above the breakdown voltage such that even a single particle (including a photon) can

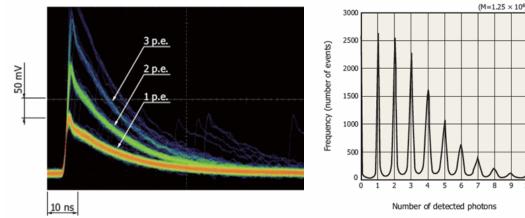
trigger an avalanche



- Provides high gain (similar to PMT)
- Excellent single photon resolution
- Non-linear output at high incident flux (saturation of pixels)
- High noise due to thermal carriers (~ 100 kHz/mm<sup>2</sup>)
- Large temperature dependence ( $\Delta$ G/G ~ 2%/°C)
- Insensitive to magnetic fields
- Very susceptible to neutron damage



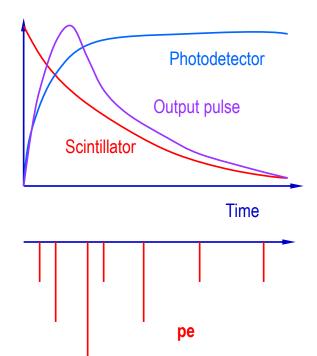




# **Microchannel Plates (MCP)**

A MCP is a thin glass plate with an array of holes forming a continuous dynode structure. Since the avalanche process is highly contained and the transit time spread is very small, these devices provide very fast timing and excellent time resolution.

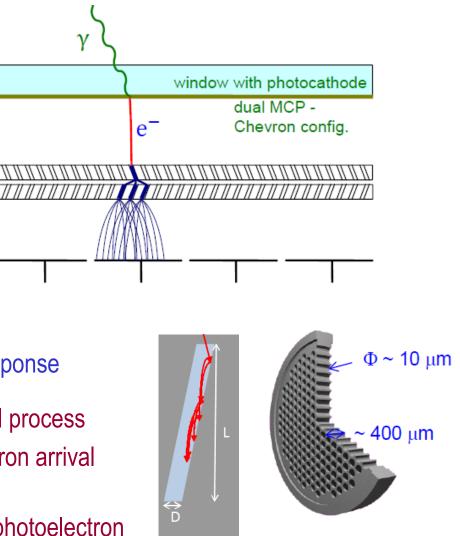
### **Timing Resolution:**



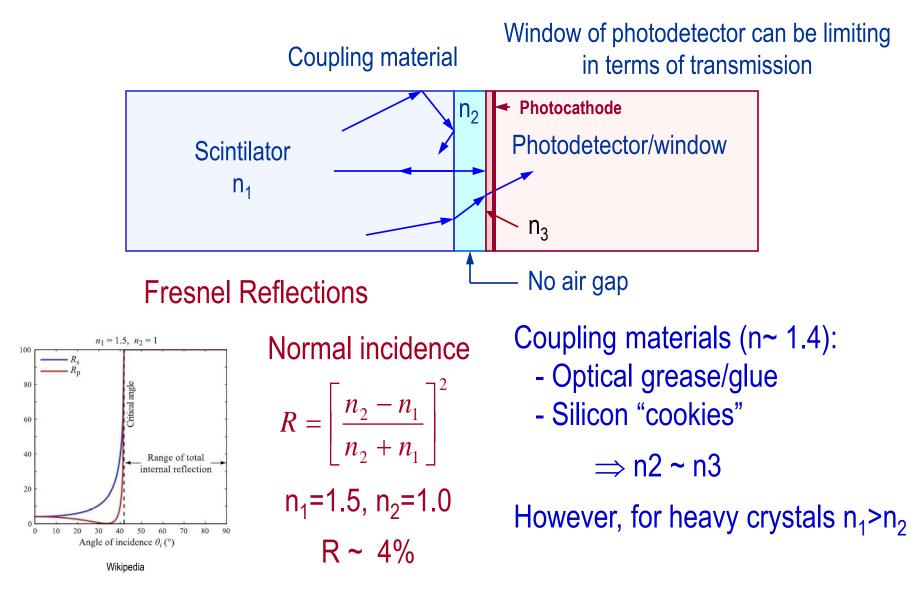
Time resolution determined by:

- Decay time of detector material
- Time response of photodetector
- Light yield of detector material
- Cherenkov materials give fastest time response
  - Photoelectron arrival time is a statistical process If light yield is low, individual photoelectron arrival times will be seen

Best timing achieved by detecting first photoelectron

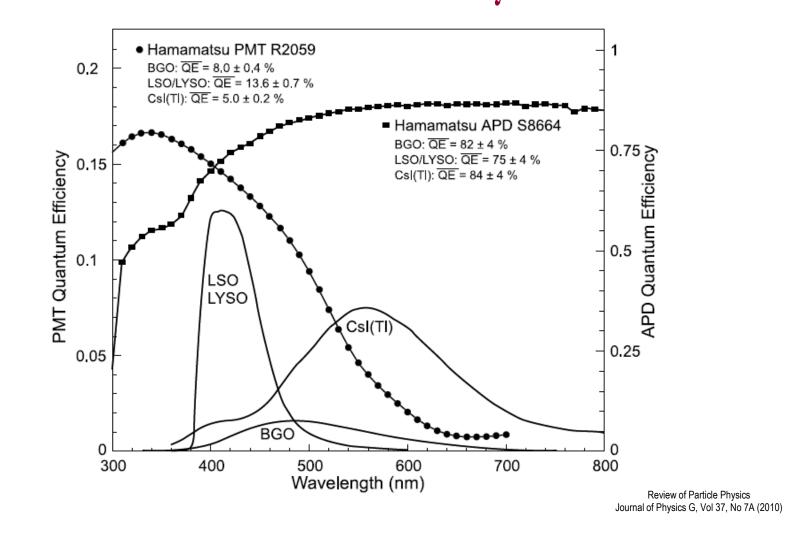


## **Coupling to Photodetector**

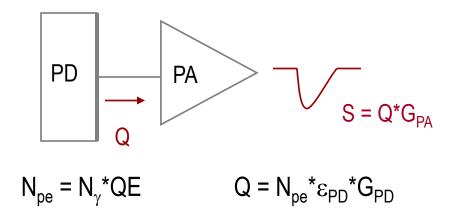


### Matching the Photodetector to the Scintillator

Emission Weighted Quantum Efficiency:  $\overline{Q} \overline{E} = \int Q E_{PD}(\lambda) \cdot I_{em}(\lambda) d\lambda$ 



### **Photoelectrons and Noise**



Photoelectrons are the primary signal and can never be recovered at a later stage in signal processing

Figure of Merit : N<sub>pe</sub>/MeV

**Energy Resolution** 

$$\frac{\sigma(E)}{\langle E \rangle} = \sqrt{\frac{ENF}{N_{pe} \cdot \varepsilon_{PD}} + \left(\frac{ENC}{N_{pe} \cdot \varepsilon_{PD} \cdot G_{PD}}\right)^2}$$

ENF = Excess noise factor from photodetector

ENC = Equivalent Noise Charge in electronic gain stage

Low noise amplification at the early stage of signal processing is crucial for achieving good signal to noise

### **Excess Noise Factor**

### **Multiplication fluctuations are characterized by the Excess Noise Factor - ENF**

- If photons are Poisson distributed so are photoelectrons with average  $\overline{n}_{pe} = PDE \cdot \overline{n}_{\gamma}$
- After multiplication with average gain and variance  $M, \sigma_M^2$  we get average output signal  $\overline{n} = M \cdot \overline{n}_{pe}$  and

$$\frac{\sigma_n^2}{\overline{n}^2} = \left(1 + \frac{\sigma_M^2}{M^2}\right) \cdot \frac{\sigma_{pe}^2}{\overline{n}_{pe}^2} = ENF \cdot \frac{1}{\overline{n}_{pe}} = ENF \cdot \frac{1}{\overline{n}_{PDE}} \cdot \frac{1}{\overline{n}_{\gamma}}$$

• Excess noise factor is

$$ENF = \frac{\sigma_{out}^2/N_{out}^2}{\sigma_{pe}^2/N_{pe}^2} = 1 + \frac{\sigma_M^2}{M^2} = \frac{\langle M^2 \rangle}{\langle M \rangle^2}$$

• Impact on photon counting capability and energy resolution

$$\frac{\sigma_E}{E} = \sqrt{\frac{ENF}{PDE}} \sqrt{\frac{1}{N_{\gamma}}}$$

Sensor	ENF
PMT	1-1.5
APD(Si)	~3 @ gain=50
HPD, HAPD	~1
SiPM	1-1.5
MCP-PMT	1-1.5

## **Applications of Scintillation Detectors**

### Nuclear and Particle Physics

- Calorimetry
- Tracking
- Particle ID (TOF)
- □ Astrophysics
  - Dark Matter
- Medical Imaging
  - PET
  - SPECT
- Homeland Security
  - Nuclear Non-Proliferation
  - Nuclear Waste Management

□ And many more...

## Calorimetry

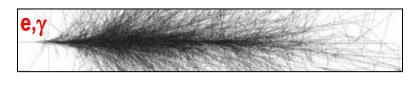
The purpose of calorimetry in particle physics is to measure the total energy of high energy particles.

Energy comes in two forms: electromagnetic and hadronic

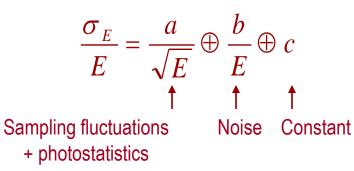
Calorimeters come in two types: sampling and homogenous

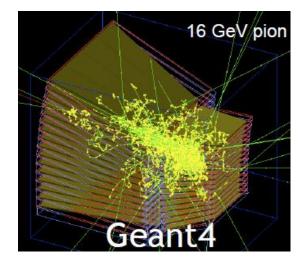
Requirements for good energy resolution:

- Detect as much of the total energy as possible (best for homogenous)
- Equalize response for electromagnetic and hadronic energy (e/h $\rightarrow$ 1)
- High light yield ( $\Rightarrow$  good photostatistics)
- Good uniformity (both longitudinally and transverse)



Electromagnetic

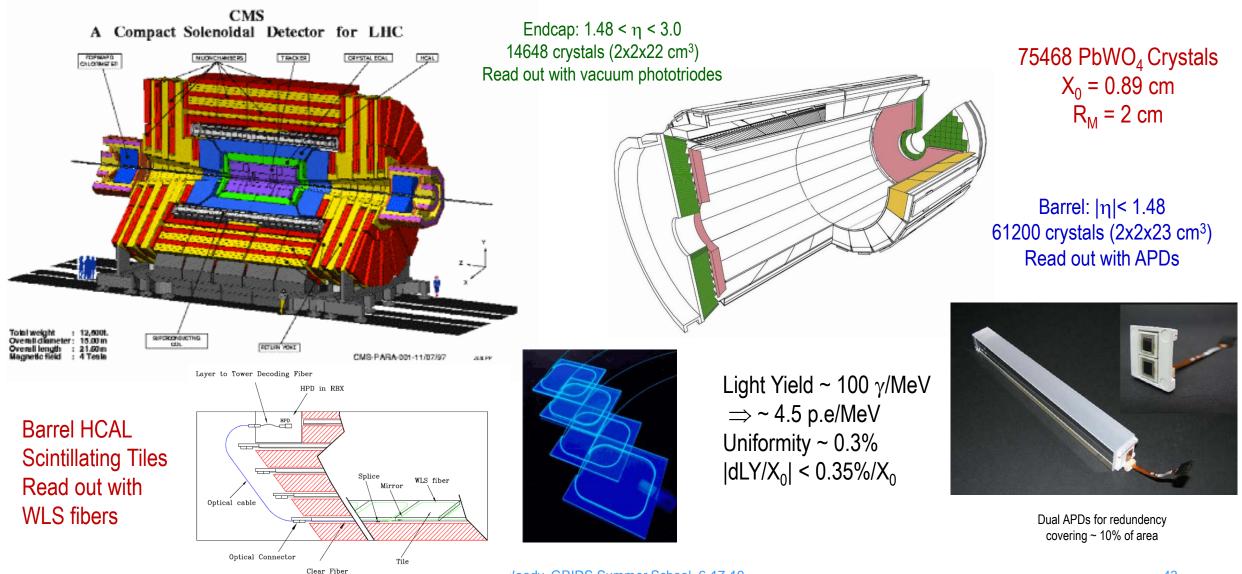




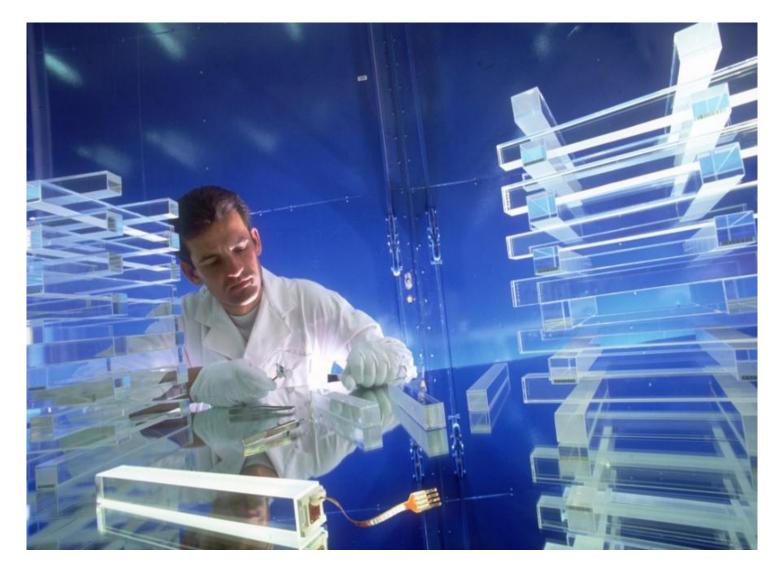
Hadronic

## **CMS** at the **CERN LHC**

Uses both sampling (HCAL) and homogeneous (EMCAL) calorimeters

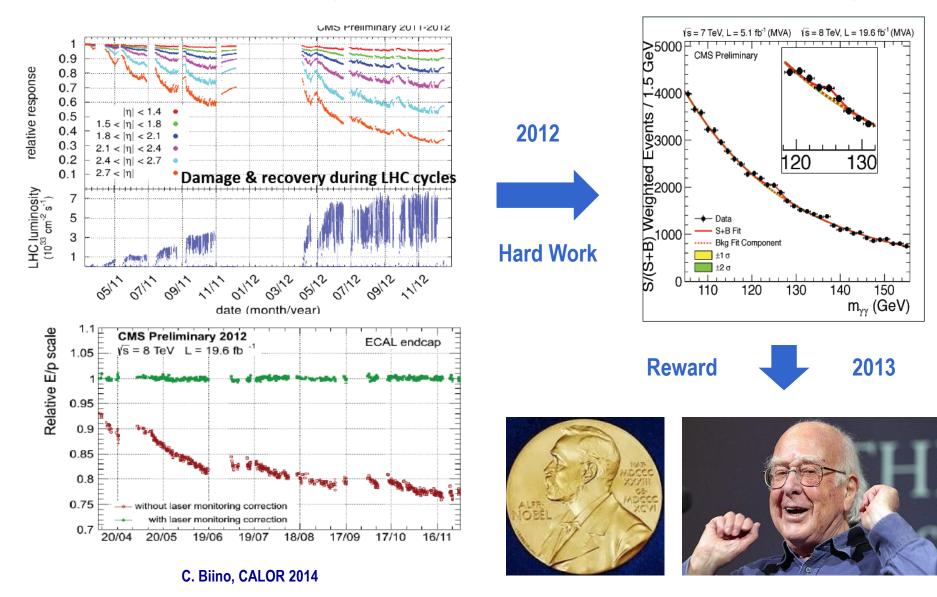


## **CMS Lead Tungstate Crystals**

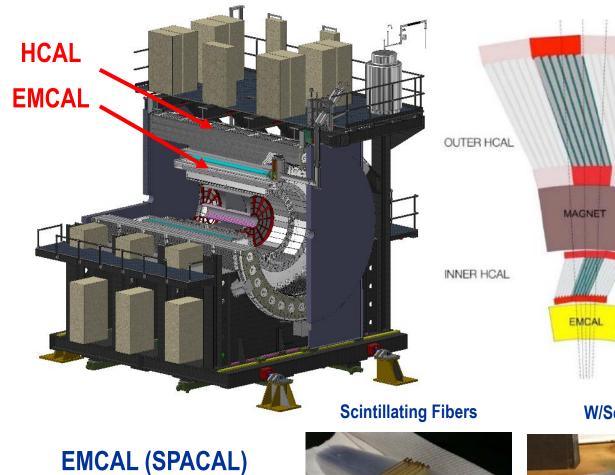


### Crystals produced in Russia and China

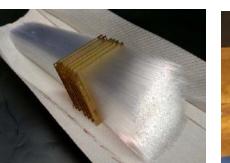
### **Radiation Damage in PWO and Response Monitoring**



## **The sPHENIX Calorimeters**



Absorber: Tungsten Powder/Scintillating Fiber/Epoxy Matrix







Scintillating Tiles Read Out with WLS Fibers





#### Fibers Ends Read Out With SiPMs

Readout end of Absorber Block

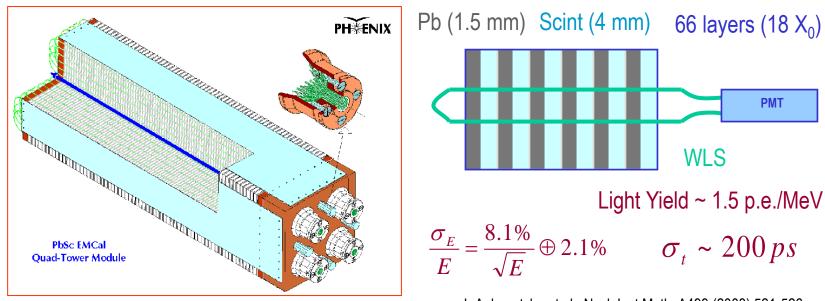


#### 4 Towers Read Out With SiPMs



## Shashlik (Shish-Kebab) Calorimeters

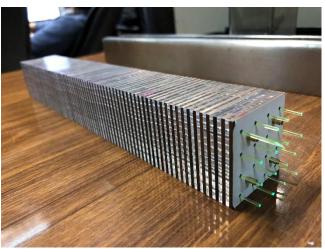
Alternating stack of lead and scintillator plates Wavelength shifting fibers pass longitudinally through the stack Fibers are bundled in the back and read out with PMTs



L.Aphecetche et.al., Nucl. Inst Meth. A499 (2003) 521-526

EIC Calorimeter R&D W/Cu Plate Shashlik





Read out individual WLS fibers with SiPMs

## **Search for Dark Matter**

# Possible Candidate: Weakly Interacting Massive Particles (WIMPs) interact only through weak interaction (like neutrinos)

WIMP interacts with nucleus, producing nuclear recoil

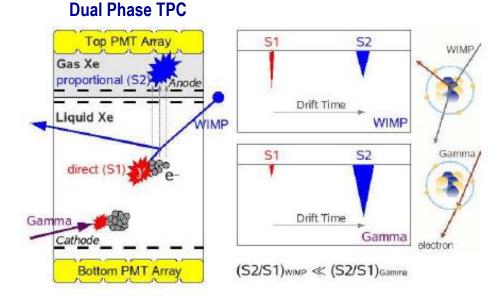
- Large direct scintillation signal (S1) in LXe, small ionization charge
- Gammas (background) interact with electrons
  - Small direct scintillation signal, large ionization charge

Ionization charge drifts to gas phase Xe where it is extracted with a large electric field and amplified in proportional mode giving second delayed scintillation signal (S2)

Discriminate WIMPs from gammas by measuring ratio S2/S1

E. Aprile, Rev. Mod. Phys, Vol 82 (2010)

Prompt signal (S1) produced in LXe Delayed signal (S2) produced in gaseous Xe



## XENON 1T Dark Matter Experiment Underground Experiment at Grand Sasso



#### **XENON 1T Cryostat and Assembly Hall**

**PMT** 

Array

TPC

Matter Projec

## **Scintillation Detectors for PET**

(Positron Emission Tomography)

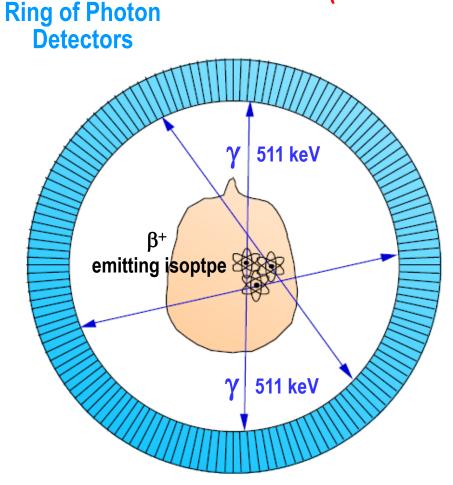
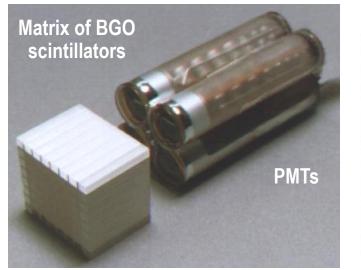
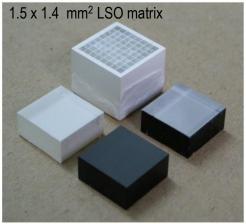


Figure courtesy of W.Moses

511 keV photons require high stopping power  $\Rightarrow$  Crystals such as BGO, LSO, LYSO ( $\rho \sim 7 \text{ g/cm}^3, \mu^{-1} \sim 1.2 \text{ cm}$ )

Light output: BGO ~ 8400  $\gamma$ /MeV LSO ~ 30,000  $\gamma$ /MeV





Light sharing (Anger Logic) PMT readout

Highly pixellated SiPM readout

## **Time of Flight PET**

### 10 ps Challenge

Achieving a Coincidence Time Resolution (CTR) < 10 ps

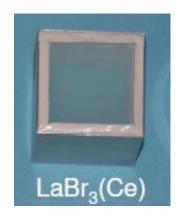
**Figures courtesy of W.Moses** 

- Improve image quality
- Reduce patient exposure
- Reduce scan time
- Enable easier whole body imaging

ithout TOF POR Annihilation  $t_2 - t_1$ Reduces noise in image True Variance reduction ~  $2D/c\Delta t$ Random  $\Delta t$  500 ps  $\Rightarrow$  x5 reduction in variance Scatter **3D PET** 

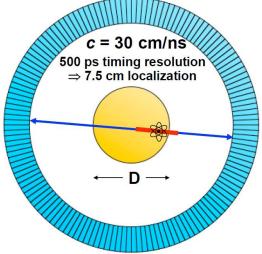
Large randoms background in PET Use timing information to localize decay point along line

Requires fast, bright crystal such as LaBr<sub>3</sub>



σ<sub>t</sub> < 100 ps σ<sub>E</sub> < 4% @ 511 keV

Also requires high speed CFD and TDC  $\Rightarrow$  ASICs





### **GE SIGNA PET/MR Scanner**

## **PET/TOF/MR Scanners**

### MRI

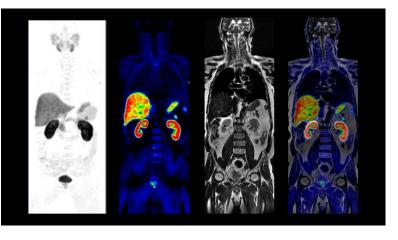
- 3T Magnet
- 50 cm radial MRI FOV

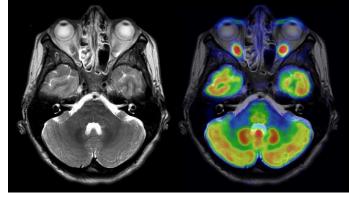
### PET

- SiPM readout
- 25 mm LBS(~ LSO) crystals
- < 400ps timing resolution
- 25cm axial FOV
- 21cps/kBq for count rate capability



**PET Detector Module** 

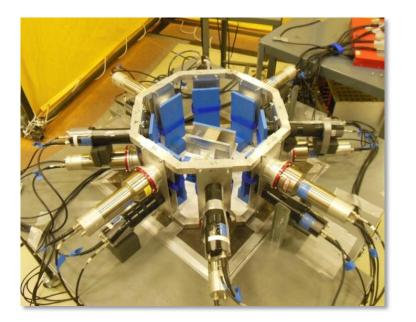




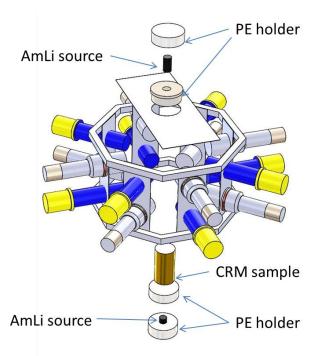
## **Nuclear Waste Management & Non-Proliferation**

#### **Passive Assay Systems**





Assay of nuclear fuel measuring n-γ spectra using liquid and crystal organic scintillators



## Measure n-γ spectra using neutron activation

#### **Detection of Nuclear Materials**





## Summary

- Scintillation detectors are widely used in many areas of particle physics, medical imaging, homeland security and many other applications.
- There exists a wide variety of scintillating materials to choose from ranging from inexpensive plastic scintillators for charged particle detection to high density crystals for detecting high energy electrons and gammas.
- There are also a wide variety of photodetectors that can be used in conjunction with many types of scintillators ranging from conventional photomultiplier tubes to many modern solid state detectors.
- A good overall detector design requires looking at all factors affecting performance (choosing the right photodetector, good light collection and matching, and low noise electronics)

## References

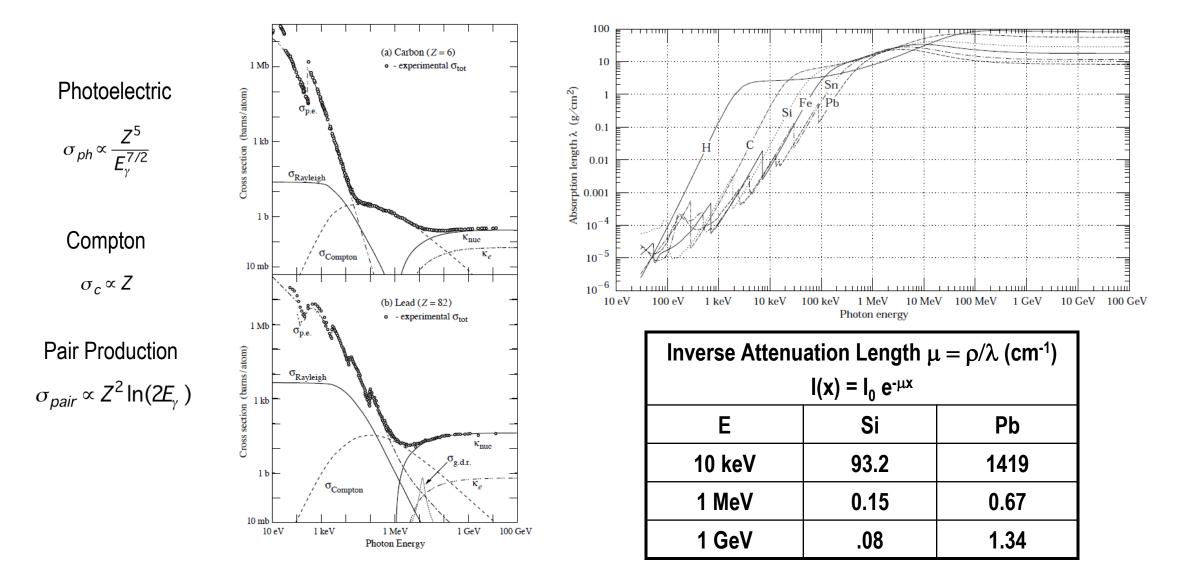
- **Radiation Detection and Measurement, Glenn F. Knoll**
- □ Inorganic Scintillators for Detector Systems, P. Lecoq, A. Gektin, M. Korzhik
- □ Calorimetry Energy Measurements in Particle Physics, R. Wigmans

Acknowledgements

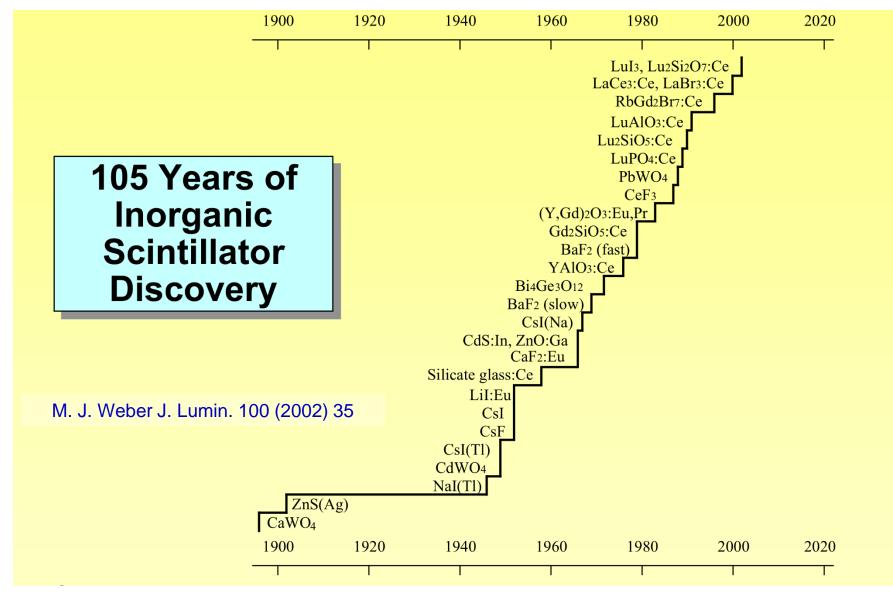
- Paul Lecoq (CERN)
- □ Bill Moses (LBNL, Retired)
- Samo Kopar (Josef Stefan Institute Ljubljana and University of Maribor)
- □ Stefan Gundacker (CERN)
- Ren Yuan Zhu (Caltech)
- □ Sara Pozzi (University of Michigan) and John Valentine (LBNL)

## **Backup Slides**

## **Photon Interactions in Matter**

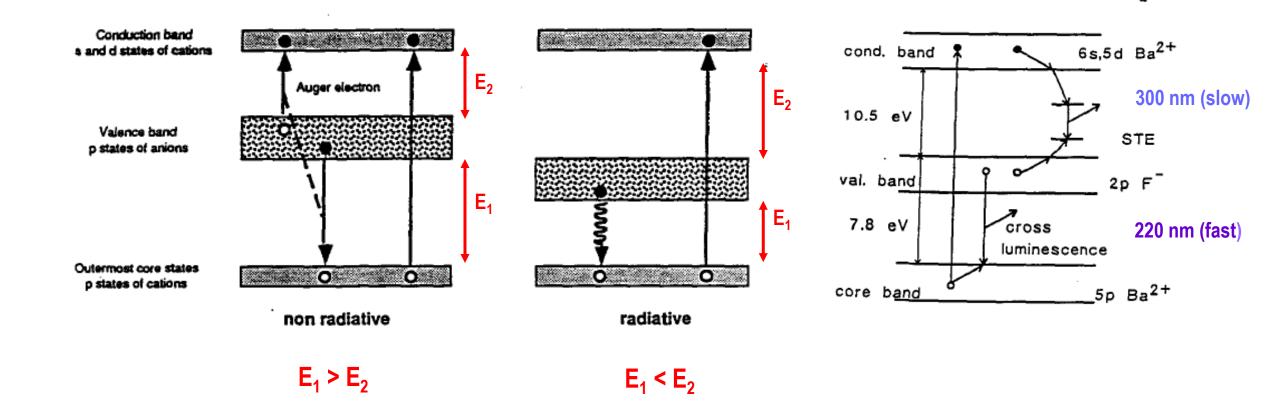


## **History of Crystal Scintillators**



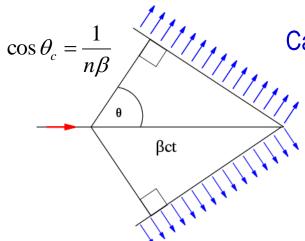
### **Cross Luminescence**

### **Example : BaF<sub>2</sub> Fast and Slow Components**



BAND STRUCTURE BaF,

## **Cherenkov Radiators**



Can be used as threshold detectors or Ring Imaging Detectors

Sensitive to most relativistic particles, so for calorimetry, measures electromagnetic component of shower

	Air	Aerogel	Water	Quartz	PbGI	PbF2
Density (g/cm <sup>3</sup> )	1.2x10 <sup>-3</sup>	0.2-0.4	1.00	2.2	6.3	7.77
Index of refraction	1.0003	1.01-1.10	1.33	1.46	1.8	1.78
Radiation Length (cm)	3.1x10 <sup>4</sup>	68-136	36	12.3	1.26	0.93
N <sub>pe</sub> /cm for N <sub>0</sub> =100	0.06	9	43	53	69	68

$$\frac{dN}{d\lambda} = 2\pi \alpha \sin^2 \theta_c \frac{1}{\lambda^2} \implies \text{Spectrum peaks in the deep (V)UV}$$

$$N_{pe} = L \frac{2\pi \alpha}{hc} \int \varepsilon(E) \sin^2 \theta_c(E) dE \qquad \varepsilon = \varepsilon_{col} \cdot \varepsilon_{det} \qquad \stackrel{n=n(E) \text{ for a dispersive medium}}{\text{dispersive medium}}$$
Figure of Merit :  $N_0 = 370 \, cm^{-1} eV^{-1} \int \varepsilon(E) dE$ 

$$N_{pe} = LN_0 \langle \sin^2 \theta_c \rangle$$

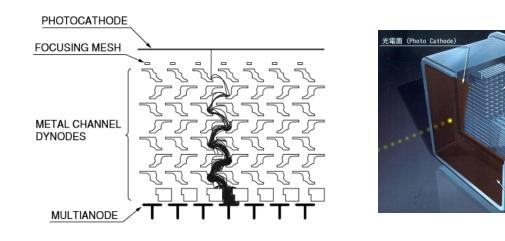
## **Multianode PMTs**

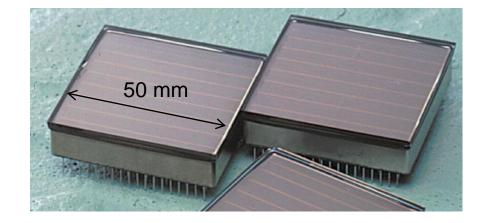
Metal channel dynode (Hamamatsu):

- multiplication is confined in a narrow channel
  - $\rightarrow$  multi-anode designs
  - $\rightarrow$  some tolerance to modest magnetic field
- ~ 30 mm x 30 mm
- gain up to 10<sup>7</sup>, excellent single photon detection
- gain uniformity typ. 1 : 2.5;
- cross-talk typ. < 2% (for 2x2 mm<sup>2</sup> pads)
- low DCR, few counts/cm<sup>2</sup>/s

Flat-panel (Hamamatsu H8500):

- . 8 x 8 channels (5.8 x 5.8 mm<sup>2</sup> each)
- ~ 50 mm x 50 mm
- Excellent active area coverage (89%)





電子增倍部 (Dynode

入射窓 (Input Window

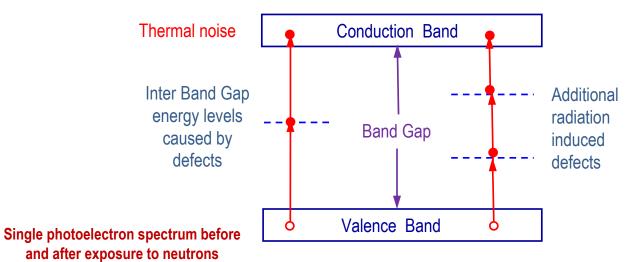
## **Radiation Damage in SiPMs**

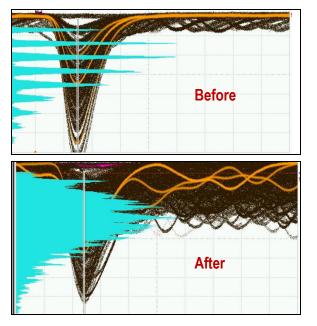
### Displacement damage

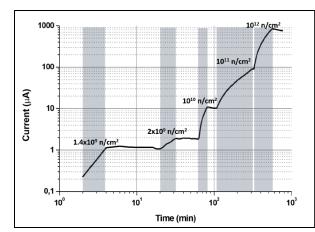
- Caused by neutrons or charged particles with energies ~ 1 MeV or higher interacting in silicon
- Knocks out a Si atom from the lattice leaving behind a vacancy and an interstitial atom, causing a defect
- Defects create energy levels in the band gap, allowing easier thermal excitation of valence band electrons to the conduction band
  - $\Rightarrow$  Higher thermal noise and increased dark current
- Noisy pixels can lead to loss in PDE

### **Ionization Damage**

- Caused by gammas or charged particles
- Can cause charge buildup in insulating materials (e.g., SiO<sub>2</sub>) which can distort the electric field in the device
- Can also cause optical damage in the window leading to transmission loss

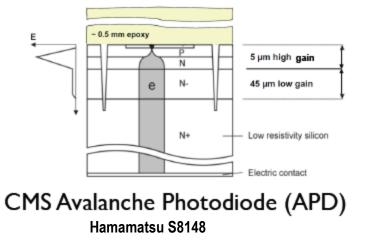


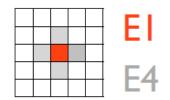




Increase in dark current after exposure to neutrons

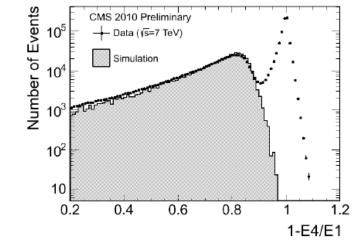
## **Nuclear Counter Effect in APDs**

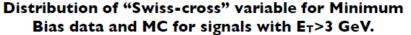




Direct ionization hits produces anomalous pattern compared with normal EM shower

- NCE signal produced by direct ionization in silicon by charged particles or neutrons
- Specially developed APD to reduce this effect
- Thin high gain depletion layer (G ~ 50,  $t_{eff}$  ~ 5 µm) followed by thicker low gain layer (G~1.4,  $t_{eff}$  ~ 45 µm)
- Expect ~ 100 MeV equivalent energy for MIP
- Equivalent neutron energy can be several hundred GeV

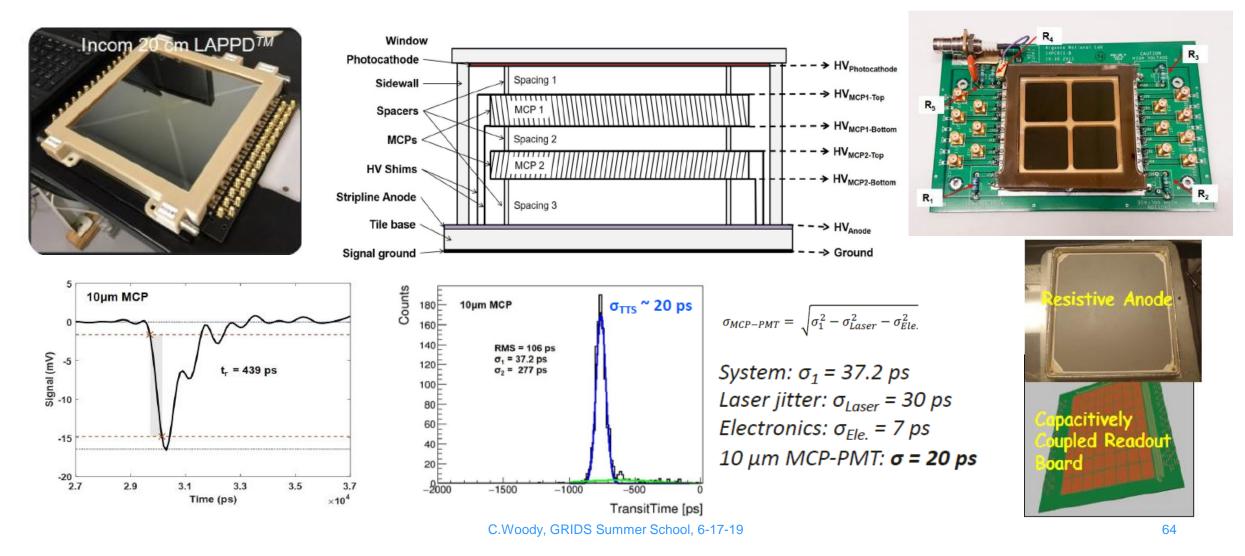




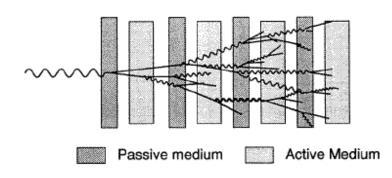
D.A.Petyt, CALOR 2012

## Large Area Picosecond Photodetectors (LAPPDs)

Large Area Pixellated Microchannel Plate detectors Developed to address the needs of the HEP and NP community



## **Sampling Calorimetry**



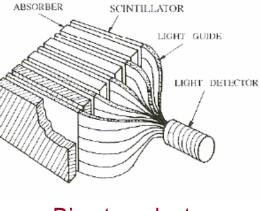
Alternating layers of passive material (absorber) and active medium to measure the energy of high energy particles

With scintillator as the active medium, problem is to collect the light onto the photodetector

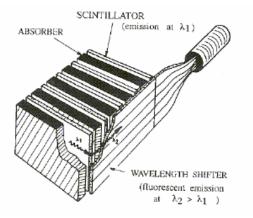
Most direct collection techniques produce unfavorable geometry for stacking modules or creating calorimeters which are hermetic (i.e., without cracks or dead material)

Want good uniformity (longitudinal and transverse) to give good energy resolution

R.Wigmans, Calorimetry Energy Measurement in Particle Physics

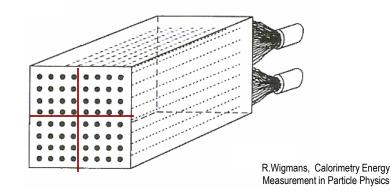


Direct readout



### WLS Bar Readout

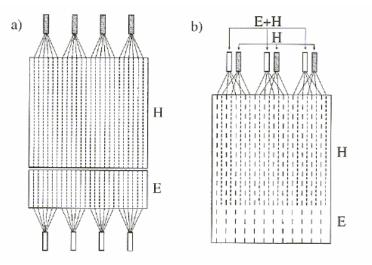
## **Fiber Calorimeters**



Fibers form the active medium which are interspersed longitudinally with the absorber material

Brings light to the front or back of the calorimeter where photodetector can be attached

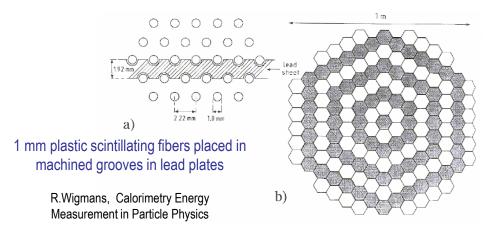
Fibers can be plastic scintillator, wavelength shifter or Cherenkov material, or a combination



Electromagnetic (front) and Hadronic (back) sections can be read out separately using different fibers

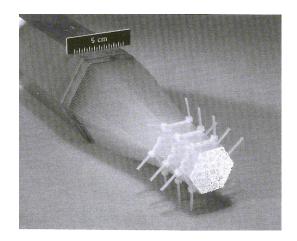
Allows almost any degree of transverse segmentation

## The SPACAL "Spaghetti" Calorimeter



First calorimeter based on scintillating fibers as the active medium

- very compact (80% absorber)
- small sampling fraction (2.3%)
- could be made hermetic with fine segmentation



**Convenient Readout** Fibers brought to back, bundled and read out with 2" PMTs

Other Sci-Fi			
Calorimeters			
H1			
KLOE			
JETSET			
CHORUS			
E864 (BNL)			

Photostatistics  $\rightarrow 300 \text{ p.e./GeV}$   $\Rightarrow \sim 6\%/\sqrt{\text{E contribution to energy resolution}}$   $\left(\frac{\sigma_E}{E}\right)_{EV} = \frac{13\%}{\sqrt{E}} \oplus 1\%$  $\left(\frac{\sigma_E}{E}\right)_{EV} = \frac{33\%}{\sqrt{E}} \oplus 2.2\%$ 

## The 'Dream Calorimeter'

DREAM = Dual REAout Module

#### **Concept**:

25

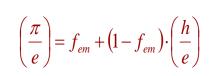
Energy resolution (%) 0 51 00

0.25

Measure electromagnetic fraction (f<sub>em</sub>) and total hadronic energy of entire shower event by event

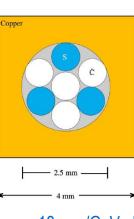
Use Cherenkov fibers to measure EM component, plastic fibers to measure (total) hadronic component

Energy resolution ultimately limited by fluctuations in nuclear binding energy loss



Copper absorber (2 m long) 4 Cherenkov fibers (quartz + clear plastic) 3 Plastic scintillating fibers

$$\frac{e}{h} \sim 1.3 \rightarrow \frac{h}{e} \sim 0.77$$

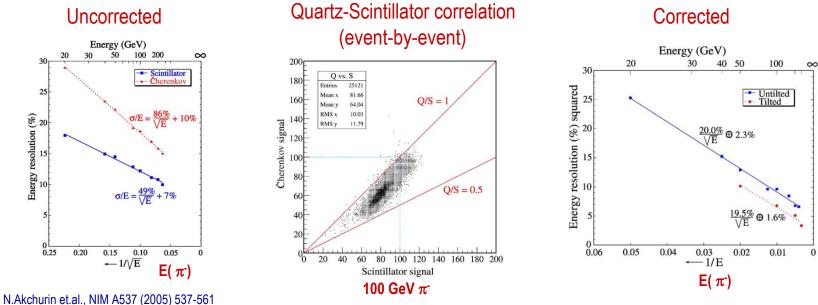


**R.Wigmans** 

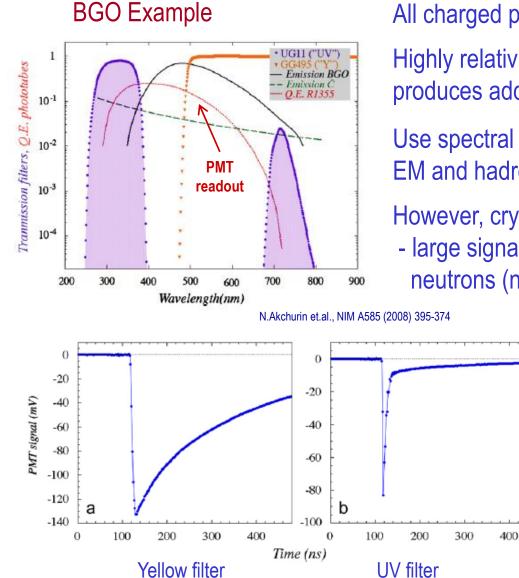
**Texas Tech University** 

~ 18 p.e./GeV plastic, ~ 8 p.e./GeV quartz

Corrected



## **Dual Readout Using Crystals**



All charged particles produce scintillation light

Highly relativistic particles (EM component of shower) produces additional Cherenkov

Use spectral and time structure differences to separate EM and hadronic components of shower

However, crystals produce large e/h

- large signal for electrons and poor sensitivity to

neutrons (not hydrogenous)

Property	Cherenkov	Scintillation
Angular Distribution	Forward peaked cone cosθ=1/nβ	Isotropic
Time structure	Instantaneous	few to hundreds of ns
Light spectrum	1/λ <sup>2</sup>	~ 400 nm

## **CMS Barrel EMCAL**



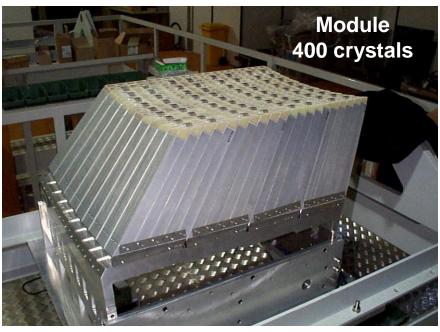
Dual APDs for redundency covering ~ 10% of area ~ 4.5 p.e/MeV Light Yield ~ 100  $\gamma$ /MeV Uniformity ~ 0.3%  $|dLY/X_0| < 0.35\%/X_0$ between 3 and 13 X<sub>0</sub>

### PbWO<sub>4</sub> crystals

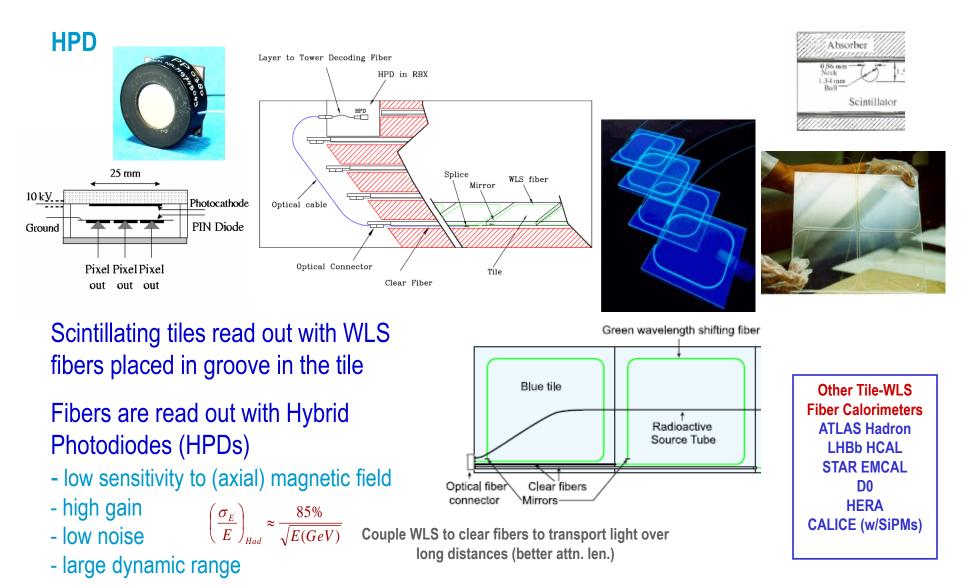
- Must be extremely uniform to obtain design energy resolution
- Must be radiation hard (achieved after extensive R&D program)

Hammamastu S-8148 (5x5 mm<sup>2</sup>)

- Insensitive to magnetic field
- Gain: M=50
- dM/dV = 3%/V, dM/dT = -2.3%/°C
- requires good V and T stability (~.05°C)
- <Q.E.> ~ 70% with  $PbWO_4$
- Nuclear counter effect (minimized depletion layer)



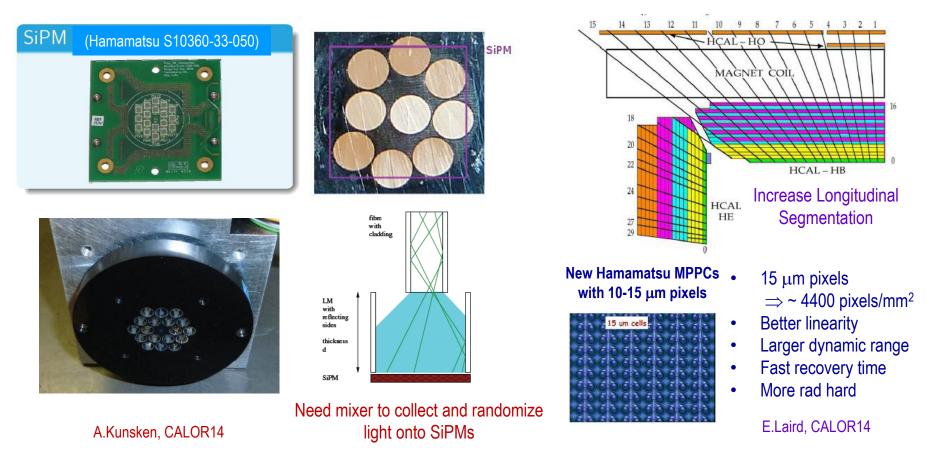
## **CMS Barrel Hadron Calorimeter**



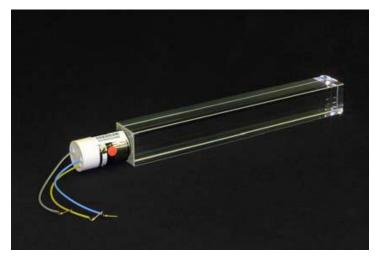
### **CMS HCAL Upgrade Plans**

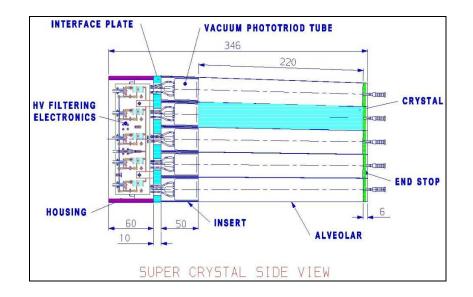
HPDs suffer from gain instability, high noise, discharging, and ion feedback Plan to replace them with SiPMs to improve performance and increase signal/noise

HO during Long Shutdown 1 (2013-2014) HB and HE during Long Shutdown 2 (2018)



## **CMS Endcap EMCAL**





~ 4.5 p.e/MeV

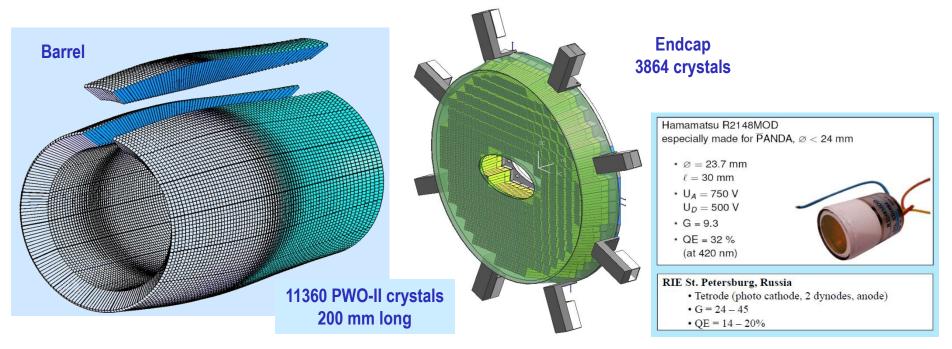
(larger area lower QE than APDs)



Vacuum Phototriode (Research Institute Electron, St. Petersburg, Russia)

- Single gain stage PMT (1" dia)
- <QE> ~ 22% at 420 nm
- Minimal sensitive to magnetic fields (15% loss at 4T)
- Gain: M=10
- Quartz window
- Rad tolerance < 10% after 20 KGy (could not use APDs due to higher radiation levels)

### **New Photosensors for PANDA**



physical goals of PANDA require further development

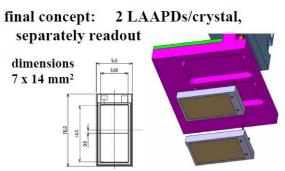
PWO-I (CMS)	PWO-II (PANDA)
420	420
100	40
150MeV - 1TeV	10MeV - 10GeV
8-12	17-22
+18	-25
3,4	2,0
	420 100 150MeV - 1TeV 8-12 +18

#### Barrel Readout Large Area APDs



10x10 mm 5x5 mm

Hamamatsu APDs

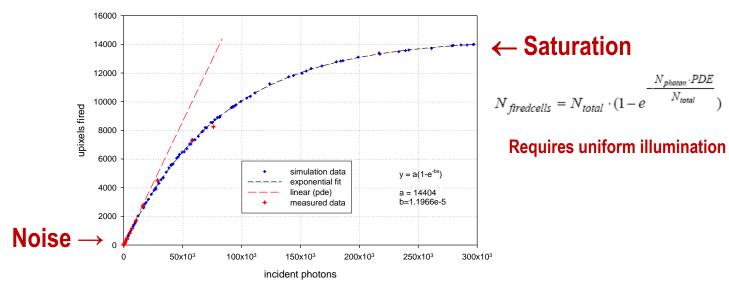


#### R. Novotny, CALOR12

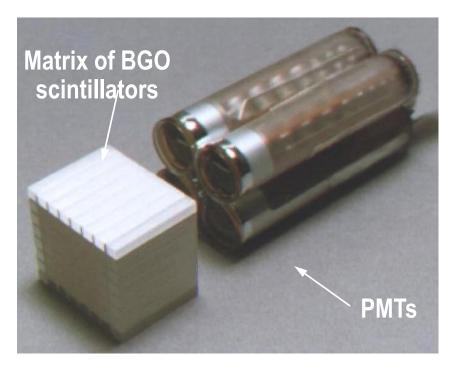
## **Use of SiPMs in Calorimetry**

Requirements				
Calorimetry	PET			
<ul> <li>Large dynamic range (~ 10<sup>3</sup>-10<sup>4</sup>)</li> </ul>	• Small device size (~ 1x1 mm <sup>2</sup> – 3x3 mm <sup>2</sup> )			
Good linearity	• Tileable			
Large area coverage	Fast timing for TOF PET			
Radiation hardness				

measured and simulated photon distribution on 3x3mm MPPC s10931-025p 14400 upixels, 25um upixels, pde=0.172 (@337nm)



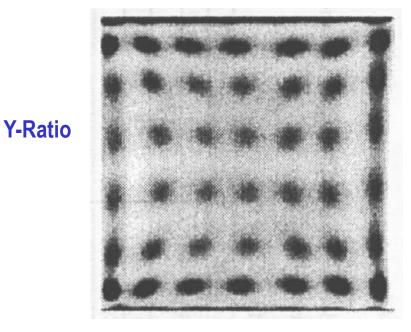
## **Standard PET Detector**



Figures courtesy of M.Casey (Siemens, formerly CTI)

PMTs and BGO have been the workhorse for PET - used in most clinical human PET scanners - new scanners are utilizing LSO/YLSO instead of BGO Position decoded by Anger Logic Spatial resolution: ~ 4 mm for human scanners ~ 2 mm for small animal scanners

Profile through Row 2



X-Ratio

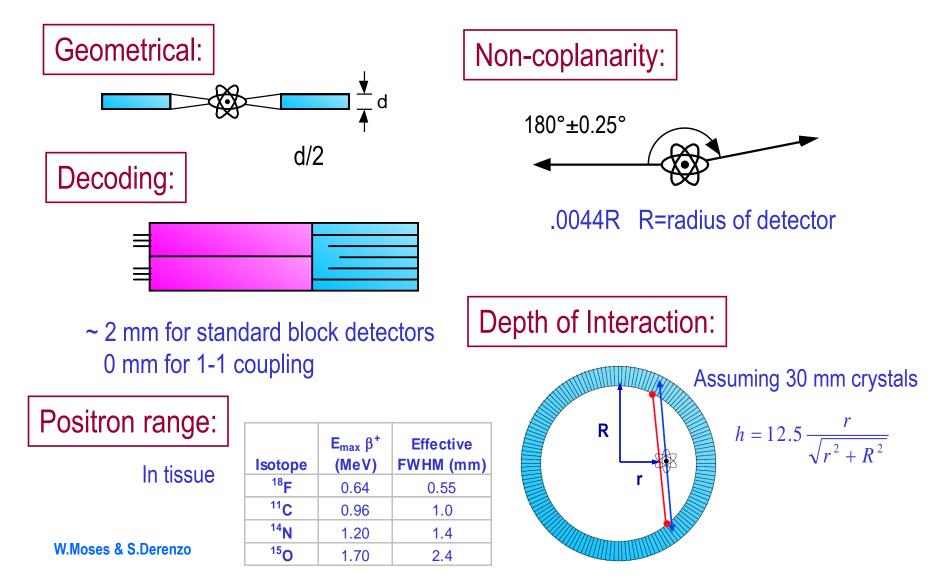
Image is "blurred" due to decoding of which pixel was hit

## **Spatial Resolution in PET**

 $FWHM = 1.25\sqrt{(d/2)^2 + b^2 + (0.0022D)^2 + r^2 + h^2}$ Geometric Decoding Non-Positron range DOI

- **1.25** : degradation due to tomographic reconstruction
- **d** : crystal size
- **b** : uncertainty due to determining block position
- **D** : coincident detector separation
- *r* : effective source size (including positron range)
- *h* : uncertainty in depth of interaction
- \* Derenzo & Moses, "Critical instrumentation issues for resolution <2mm, high sensitivity brain PET", in *Quantification of Brain Function, Tracer Kinetics & Image Analysis in Brain PET*, ed. Uemura et al, Elsevier, 1993, pp. 25-40.

## **Contributions to the Spatial Resolution**



## How Well Can One Do?

#### Center of FOV, as a function of crystal width:

#### Assumptions:

- 1-1 coupling
- <sup>18</sup>F (shortest range)
- 30 mm thickness
- 10 cm radial FOV

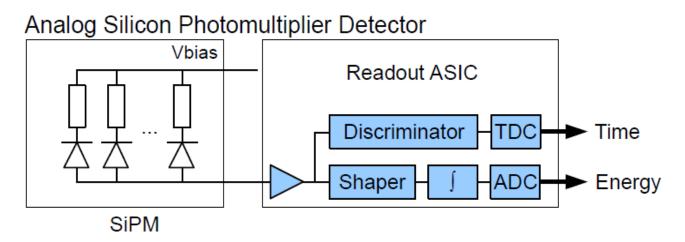
Crystal Width (mm)	1.0	0.5	0.0
Decoding (1-to-1coupling)	0	0	0
Range ( <sup>18</sup> F)	0.55	0.55	0.55
Acollinearity (R=100 mm)	0.44	0.44	0.44
Distance from Center (mm)	0	0	0
DOI (30 mm detector thickness)	0	0	0
Reconstruction (x factor)	1.25	1.25	1.25
Total (mm)	1.08	0.93	0.88

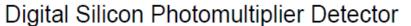
### 1 mm crystal width, as a function of distance off axis

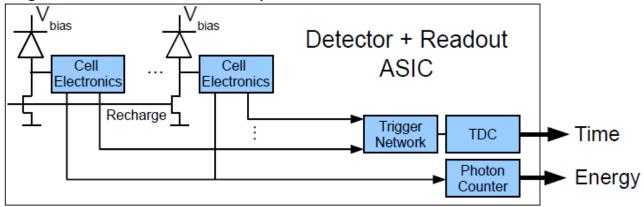
Crystal Width (mm)	1.0	1.0	1.0	1.0
Decoding (1-to-1coupling)	0	0	0	0
Range ( <sup>18</sup> F)	0.55	0.55	0.55	0.55
Acollinearity (R=100 mm)	0.44	0.44	0.44	0.44
Distance from Center (mm)	0	10	20	30
DOI (30 mm detector thickness)	0	1.24	2.45	3.59
Reconstruction (x factor)	1.25	1.25	1.25	1.25
Total (mm)	1.08	1.89	3.25	4.62

PHILIPS

### Digital SiPM – The Concept



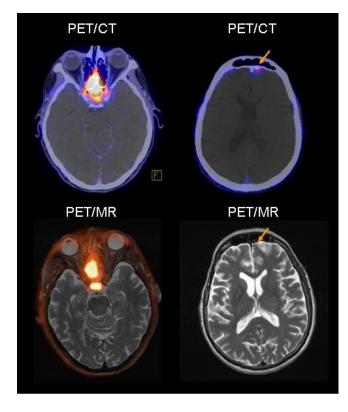




## **Simultaneous PET-MRI**

### Advantages of simultaneous PET-MRI

- MRI provides high spatial resolution for excellent anatomical information with good soft tissue contrast.
- PET provides high sensitivity and specificity for studying metabolic activities and radiotracer uptake
- PET/MR delivers ~ ½ the radiation dose of PET/CT
- Simultaneous acquisitions result in perfectly co-registered images with less scan time than sequential imaging, and also allows time correlated studies such as in combination with fMRI



B.Pichler PSMRI 2012

## **Technical Challenges with Simultaneous PET-MRI**

- PMTs used in conventional PET detectors are very sensitive to to even small magnetic fields
- MRI uses strong RF pulses at frequencies of several hundred MHz and pulsed gradient magnetic fields that can cause interference with the PET readout electronics
- PET detector and readout electronics can cause interference with very sensitive MRI detectors for measuring weak MRI signals
- Any magnetic materials in the PET detector can perturb the very homogeneous magnetic field of the MRI and can cause distortions in the MRI image